Exploring the angular dimension of ERB with the Libera camera



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- Libera camera characteristics
- Example application 1: Split-solar angular distribution model generation
- Example application 2: Testing of proposed alternate split-solar radiance-to-irradiance conversion
- Example application 3: Stereo cloud detection in challenging environments

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Sampling Schematic



• 2048 × 2048 pixel array samples entire Earth disk subtended from the satellite

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Gristey et al., AMT [2023]

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- 1.5% uniformity, 5% absolute accuracy

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 Can typically only downlink small fraction of pixel array

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- Select "ADM samples": groups of pixels encompassing Libera point spread function (PSF)
- Randomization (or "systematic shift") within angular bins from one exposure to the next
 - > 39,609/4,194,304 pixels (0.94%)



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555 nm provides optimal proxy for split-solar

 CLARREO OSSE data (*Feldman et al., JGR [2011]*) suggests midvisible wavelength is optimal for Libera VIS sub-band



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- Consistent with:
 - SCIAMACHY Gottwald and Bovensmann [2011]
 - SCIAMACHY-like simulations Gristey et al., J. Climate [2019]
 - ➢ AVIRIS Green et al., RSE [1998]
 - CERES unfiltering simulations Loeb et al., JAM [2001]



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- 555 nm has several operational advantages
 - Matches VIIRS M4 band (flat-fielding)
 - Less optical degradation



Gristey et al., AMT [2023]





2021-10-01 23:28 UTC

After 1 day of sampling...

NOAA-20 Retrieval: Minnis et al., IEEE [2021]

<u>Key</u>

• Night ADM sample

- Day ADM sample, outside VIIRS swath
- Day ADM sample, added to count

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 - 3. From the selected lookup table member, take the VIS and NIR anisotropic factors to determine the partitioning of the total shortwave irradiance into VIS and NIR irradiance:

$$\frac{F_{VIS}}{F_{NIR}} = \frac{\pi I_{VIS}/R_{VIS}}{\pi I_{NIR}/R_{NIR}} = \left(\frac{I_{VIS}}{I_{NIR}}\right) \left(\frac{R_{NIR}}{R_{VIS}}\right)$$
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		observed	retrieved
$\overline{F_{NIR}}$	$\overline{\pi I_{NIR}/R_{NIR}}$	$-\left(\overline{I_{NIR}}\right)$	$\left(\overline{R_{VIS}}\right)$
F_{VIS}	$\pi I_{VIS}/R_{VIS}$	(I_{VIS})	$\left(R_{NIR}\right)$



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F_{VIS}	$\pi I_{VIS}/R_{VIS}$	(I_{VIS})	$\left(R_{NIR}\right)$
$\overline{F_{NIR}}$	$-\frac{1}{\pi I_{NIR}/R_{NIR}}$	$-\left(\overline{I_{NIR}}\right)$	$\left(\overline{R_{VIS}}\right)$
		observed	retrieved

F : irradiance (or flux)I : radiance (or intensity)R : anisotropic factor

- One key concern: information about the split-solar angular distribution is entirely based on theory
 - > The Libera camera could provide a useful observational test here...



• Along track "stripe" - also downlinked



- Along track "stripe" also downlinked
- In the 16 minutes to pass through the camera WFOV, an exposure every 5 seconds provides <u>192 angular radiances</u>



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Cross-track imagery provides scene context



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 - Provides scene context for radiometer footprint

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- Cross-track and fwd "stripes" also downlinked
 - Provides scene context for radiometer footprint
- Camera cloud detection sometimes challenging (multispectral imagers too)
 - Cloud parallax: the apparent horizontal shift of a cloud relative to the surface with view-angle *e.g., Zhao & Di Girolamo, JAM [2004]*



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3D radiative transfer: EaR³T *Chen et al., AMT* [2023]



3D radiative transfer: EaR³T *Chen et al., AMT* [2023]



Monochromatic 555 nm image → R @ 30° fwd

Monochromatic 555 nm image → G @ nadir

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Credit: K. S. Schmidt & K. Dong

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Summary and conclusions



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555 nm

Reflectance

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"Hyper-angular" data could serve as a powerful observational test of an alternate radiance-to-irradiance approach



Stereo imagery shows promise for assisting cloud detection in challenging environments

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Counts: SZA dependence



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Counts: Cloud optical depth dependence



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Spectral relationship vs. angle



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extra

Spectral relationship: SCIAMACHY/AVIRIS

extra



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Expanding beyond along-track



extra

Expanding beyond along-track



extra



Angular index as an additional dimension



Angular index: $\alpha = (fwd - bwd) / nadir$

extra

Reflectance (standard retrieval)

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Example of cloud over frozen surface

Same cloud scene over a frozen surface



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