

Clouds and the Earth's Radiant Energy System

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**Experiments using CERES observations of the moon to
supplement the Cal/Val protocol**

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CERES Instrument Working Group**

**CERES Science Team Meeting
Virtual
October 13, 2021**

Primary CERES Flight Radiometric Validation Protocol Activities

		Product Level	Spatial Scale	Temporal Scale	Spectral Weighting	Dynamic Range	Latitude Range	Radiometric Metric	Channel	Product
On-Board	Internal BB	Filtered Radiance	Full IFOV	Continuous Capability	290-320 K BB	Across	All	Absolute Accuracy, Stability	TOT, WN, LW	-
	Internal Lamp	Filtered Radiance	Full IFOV	Continuous Capability	17000, 2000, 2300 K BB	Across	All	Absolute Stability	SW	-
	Solar	Filtered Radiance	Full IFOV	1 per orbit capability	Solar	Fixed, High	N or S Pole	Relative Stability	TOT, SW	-
Vicarious	Theoretical Line-by-Line	Filtered Radiance	>20 Km	Instantaneous	Various Earth	Across	N/A	Inter-Channel Theoretical Agreement	TOT, WN, LW	-
	Unfiltering Algorithm Theoretical Validation	N/A	N/A	N/A	N/A	N/A	N/A	N/A	TOT, SW, WN, LW	-
	Inter-Satellite (Direct Comparison)	Unfiltered Radiance	1-deg Grid	1 per crossing	Various Earth	Mid	All	Inter-Instrument Agreement, Stability	TOT, SW, WN, LW	OLR, RS
	Tropical Mean Geographical Average)	Unfiltered Radiance	20N-20S	Monthly	Tropical Ocean, All Sky	Mid	20N-20S	Inter-Instrument Agreement, Stability	TOT, WN	OLR
	DCC Albedo	Unfiltered Radiance	> 40 Km	Monthly	Cloud RS	High	All, Daytime	Inter-Instrument Agreement, Stability	SW	RS
	3-Channel Intercomparison	Unfiltered Radiance	>100 Km	Monthly	Various Earth	Across	All, Daytime	Inter-Channel consistency, Stability	TOT, SW, WN, LW	OLR, RS
	Time Space Averaging	Fluxes	Regional, Zonal, Global	Monthly, Yearly	Various Earth	Across	All	Inter-Instrument Agreement		OLR, RS
	Lunar Radiance Measurements	Filtered Radiance	Sub IFOV (7-10%)	1 day per lunar month	Lunar OLR and RS	Fixed, Low	N or S Pole	Relative Stability	TOT, SW, WN, LW	

At the end of the presentation we will compare results of the Internal Calibration Module and Lunar Observations

Executive Summary

The moon is NOT an adequate standalone calibration target for Spectrally broadband, Wide Angular Field of View, Absolute Radiometry sensors such as CERES.

However... It does play a role in the overall CERES Post Launch Cal/Val protocol

It presents itself as a highly unstable reflector (spatially nonuniform with insufficient spectral knowledge) calibration target that 'wobbles' in a galactic 'optical bench' where the key distances and illumination angles between source, reflector and measurement system vary rapidly and out of phase.

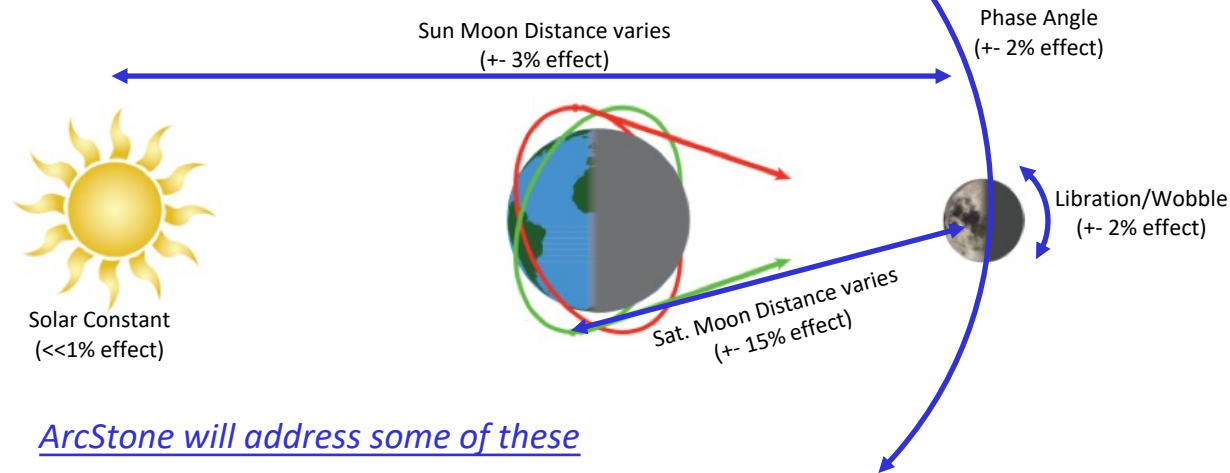
However... ROLO has, and ArcStone will in the future make important advances in characterizing these instabilities in the coming decade

Fully interpreted/corrected Lunar observations are consistent with all other elements of the cal/val protocol

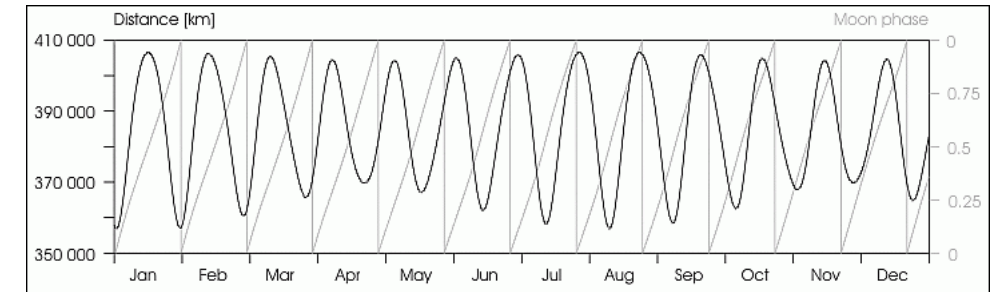
- 1) Pointing Knowledge (supplements coastline detection experiment)**
- 2) Co-alignment of the three science channels**
- 3) Spatial/Angular uniformity of dynamic response in the sensor assemblies**
- 4) Decadal trends consistent with those measured by onboard calibration sources (blackbodies and lamps)**

Lunar Cal Phenomena to be Untangled

Variation of illumination path lengths and angles



Variation of Illumination Path Lengths and Angles all Have Unique Periodicity



The Sidereal and Synodic Months. The sidereal month is the time the Moon takes to complete one full revolution around the Earth with respect to the background stars. ... Thus, the synodic month, or lunar month, is longer than the sidereal month. **A sidereal month lasts 27.322 days**, while a synodic month lasts 29.531 days.

Mismatch of Lunar Disk Target Size to Optical system

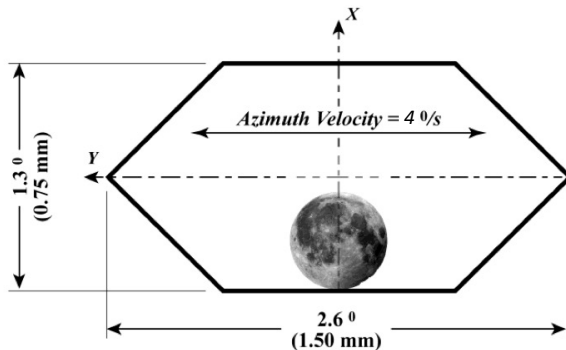
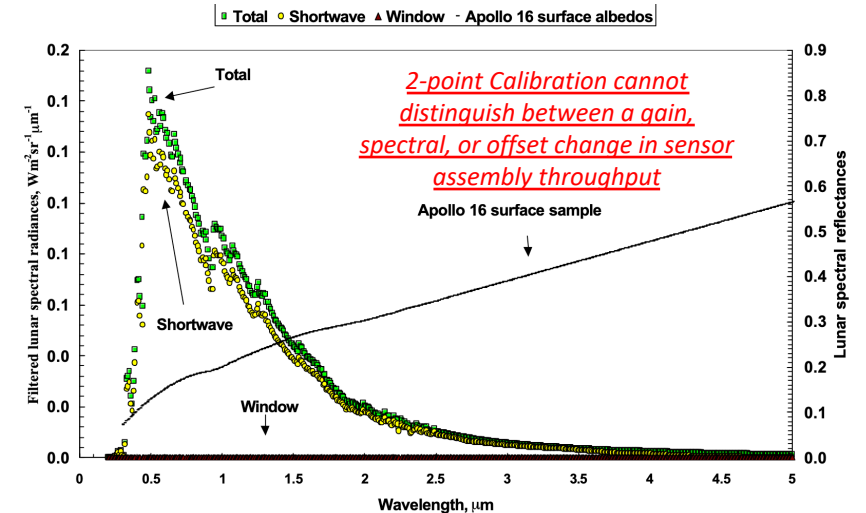


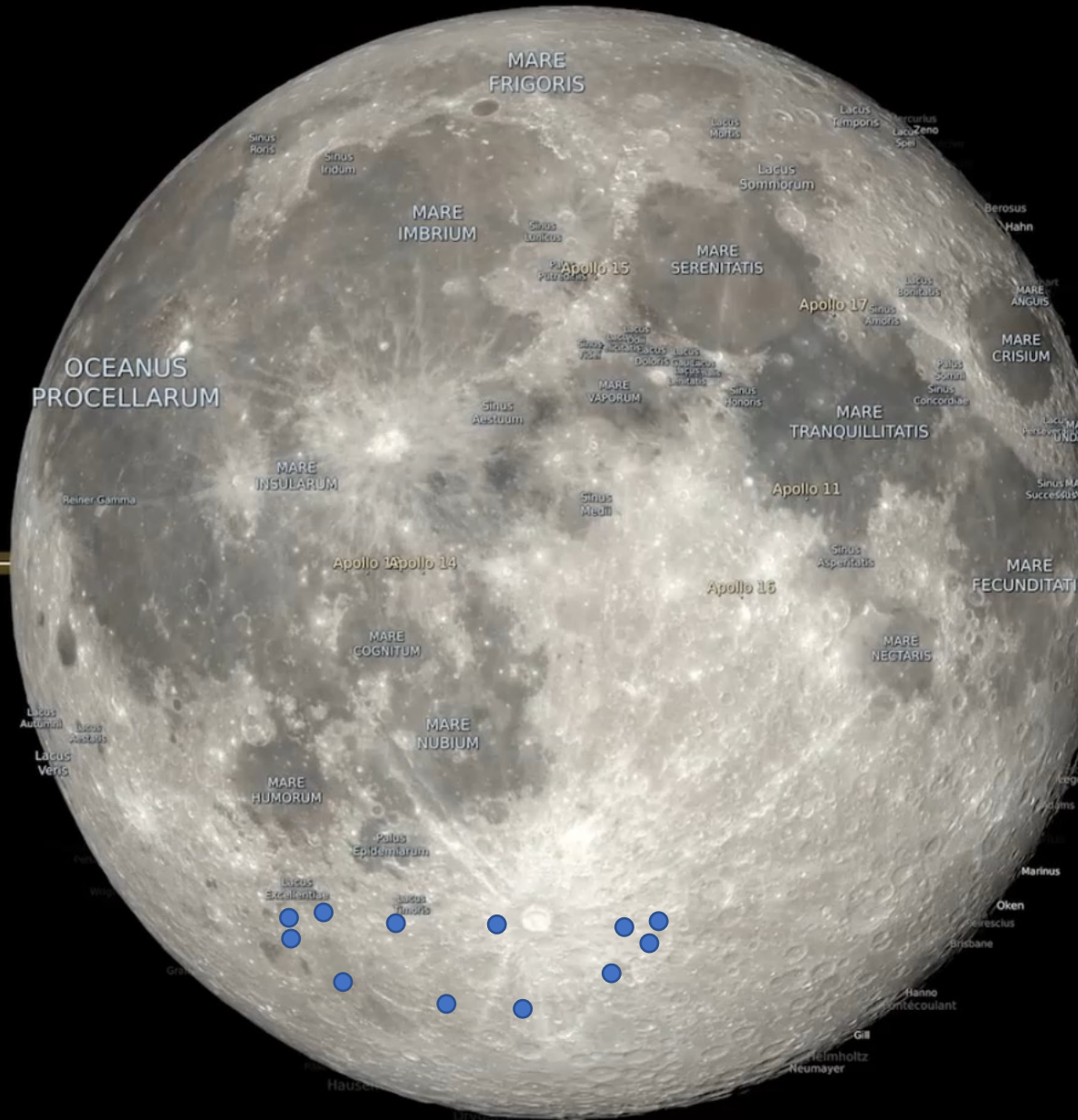
Fig. 3. Moon in CERES field of view

- Moon fills 7.4% of CERES FOV at perigee and 9.7% at apogee
 - yields time varying Solid angle of source illumination.
- >90% System etendue is 2K deep space background.
 - Yields Very low signal

Presents Minimal Variation in Integrated Spectral Radiance



~ A Case Study in how not to design a calibration experiment ~



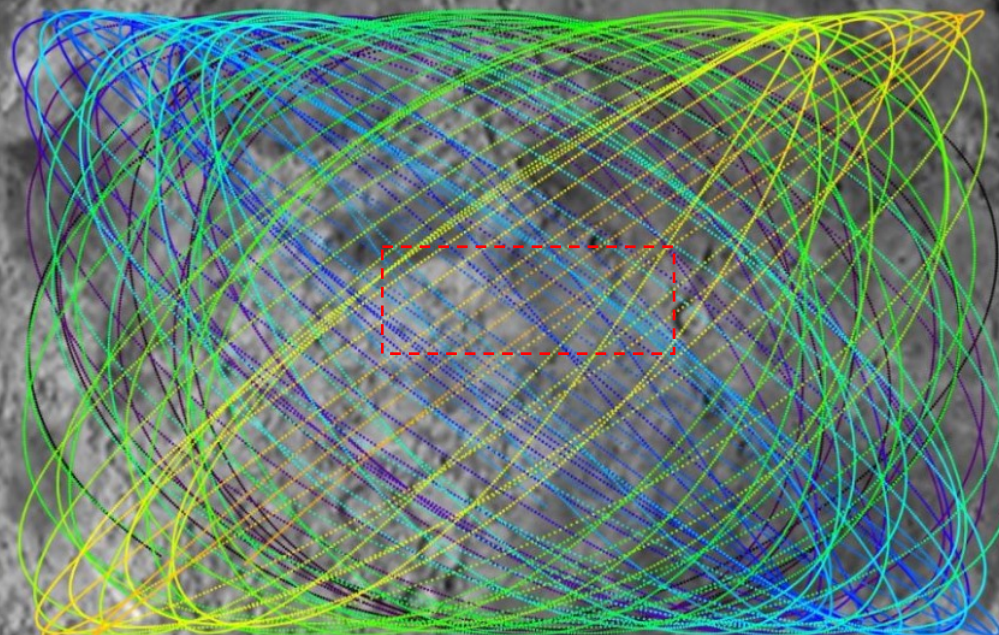
Time	20 Dec 2021 04:00 UT
Phase	99.1% (15d 20h 17m)
Diameter	1769.9 arcseconds
Distance	404963 km (31.78 Earths)
Position	06h 39m 03s, 26° 19' 27"N
Subsolar	0.753°S 12.843°W
Sub-Earth	4.168°S 2.356°W
Pos. Angle	5.096°

Moon Phases 2021

Including Libration and Position Angle



Bobblehead Effect



Note: Not To Scale



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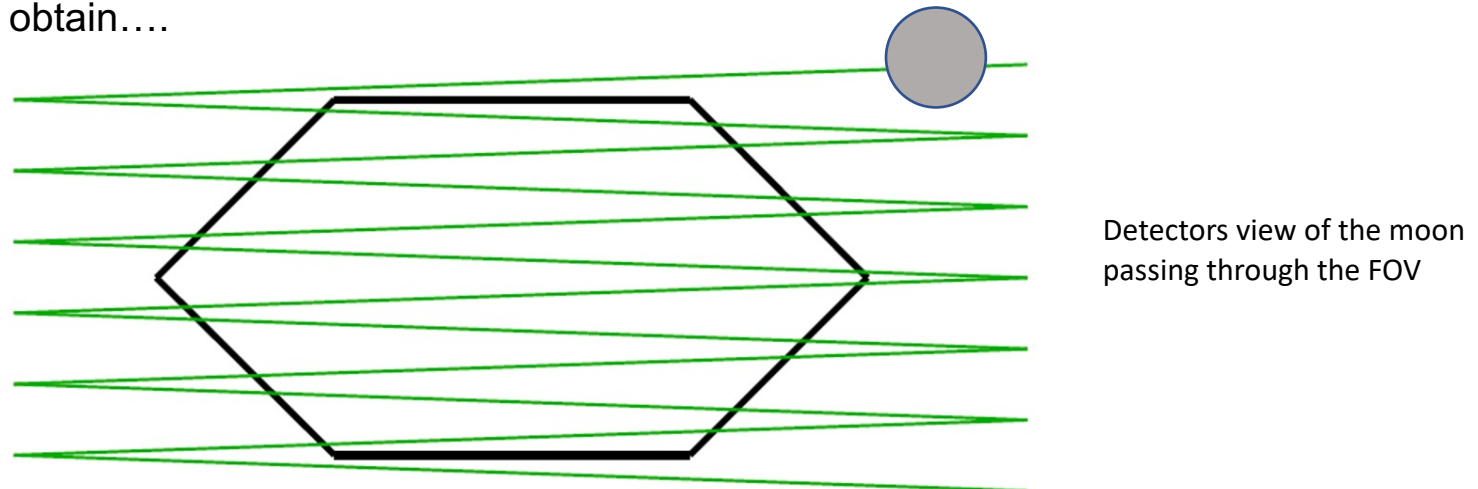
Lunar Observation Experimental Design

Objective: Utilize the moon as a quasi-point source to complete a near steady-state raster scan across the CERES FOV to support Cal/Val activities

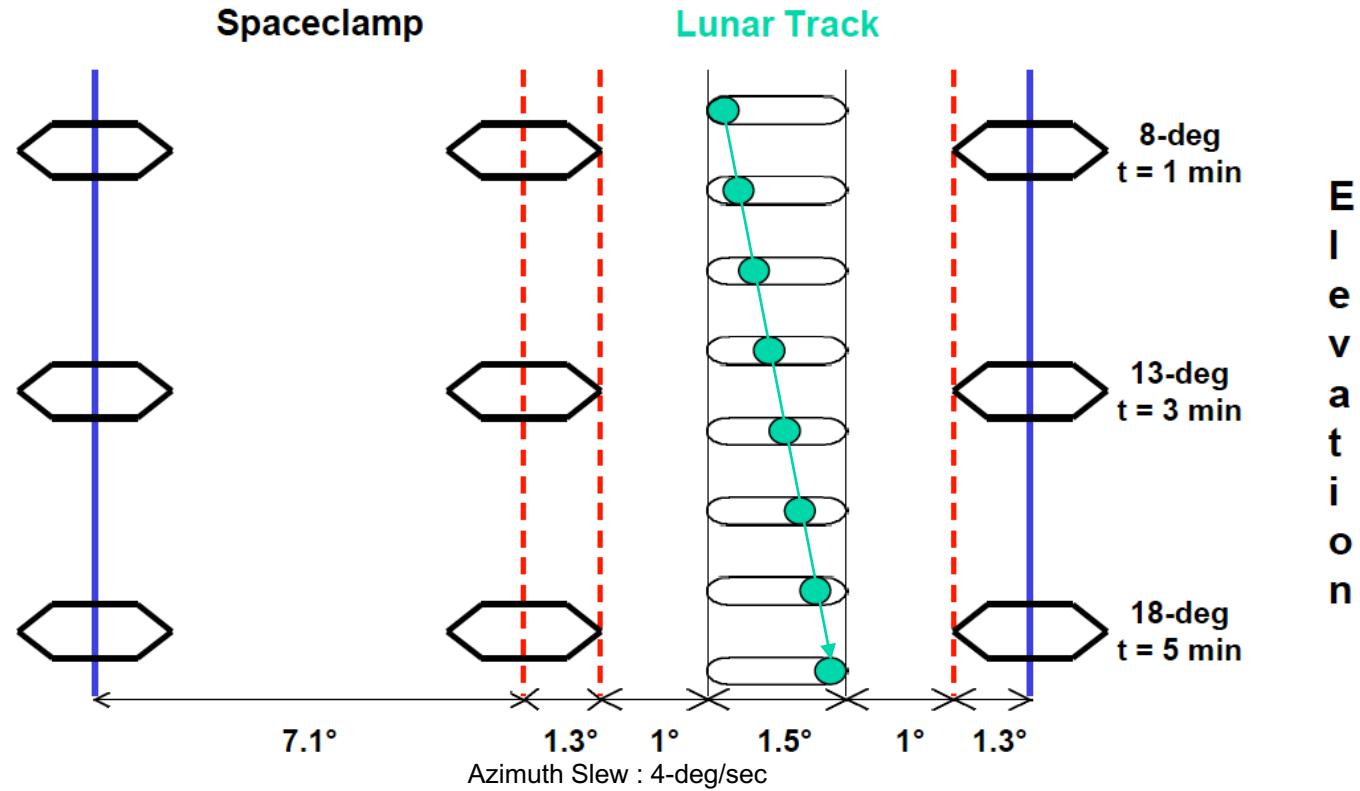
Goals

- Validate pre-launch alignment measurements (**dedicated observations began in April 2001**) (*Ipod released May 2001*)
- Measure inter-channel relative pointing accuracy, (i.e. co-registration of telescopes)
- Map out system level spatial non-uniformities in the CERES Optics/Detectors
 - Detector Requirement is $\pm 10\%$ from raster scan focused spot across the detector (component test)
 - This type of mapping is not performed **under vacuum** prior to launch
- Develop data set for long term stability measurements utilizing lunar radiances (**stable protocol since 3/2006**)

By combining knowledge of the motion of the moon relative to the spacecraft and the programmability of the CERES Instruments we obtain....



Lunar Observation Experimental Design



Elevation Gimbal is fixed at one of three pre-determined angles above the earth's limb

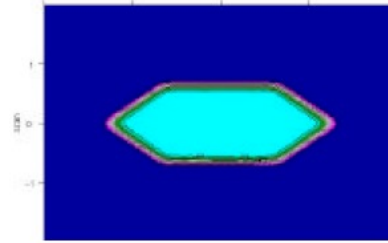
Elevation Gimbal completes an out and back rotation in 6.6 second data packet

Earth Limb

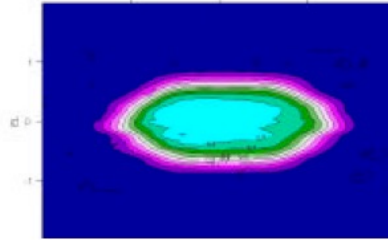
Operation is completed for the five orbits before and after full moon each lunar month

Lunar Scanning Results – CERES Optical Transfer Function

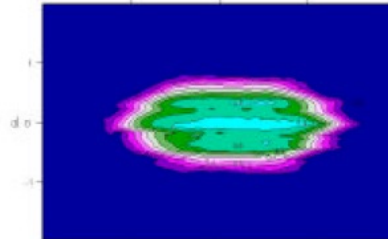
Monte-Carlo Ray Trace
(FELIX)



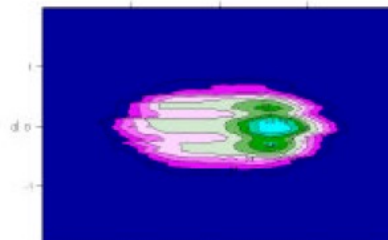
FM-2 Total



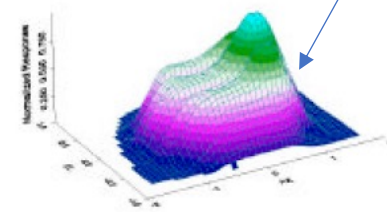
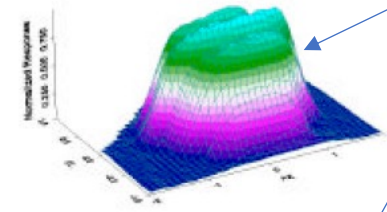
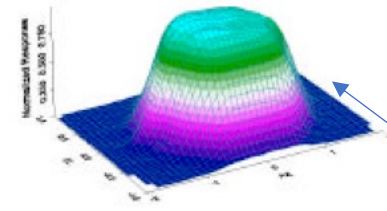
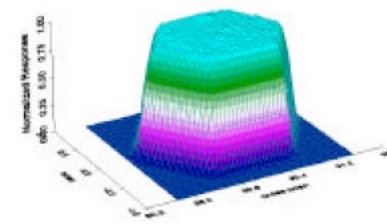
FM-2 Shortwave



FM-2 Window



Confirms alignment and
co-registration



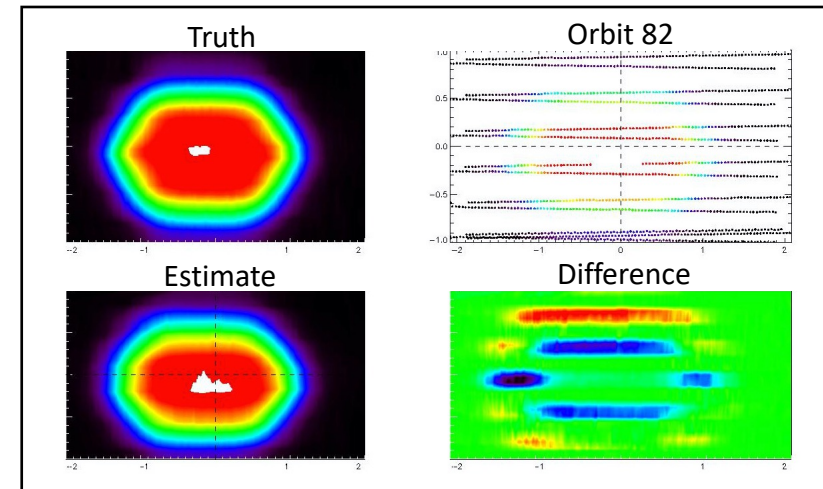
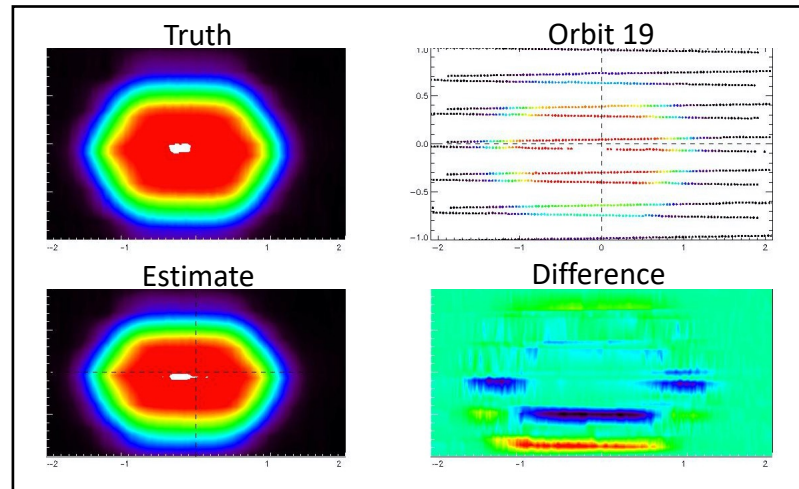
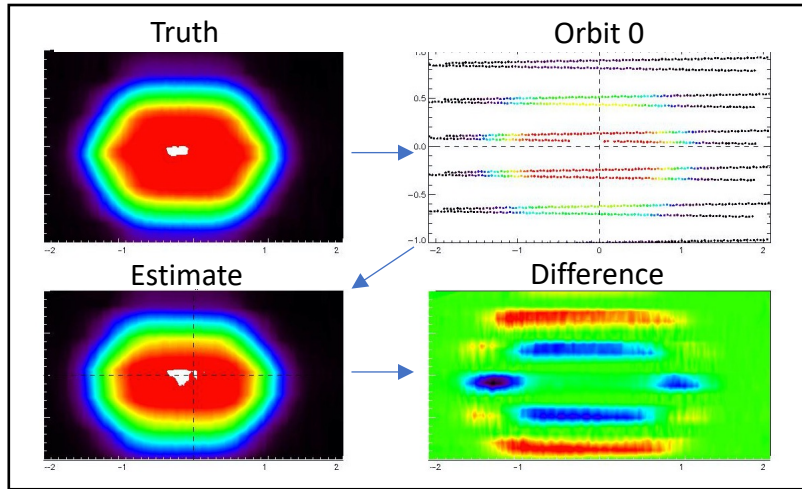
Integrating across the
transfer function
provides an estimate of
full-field lunar signal

How accurately can this be
accomplished with
sparsely sampled data?

Displays non-uniformity
in Response

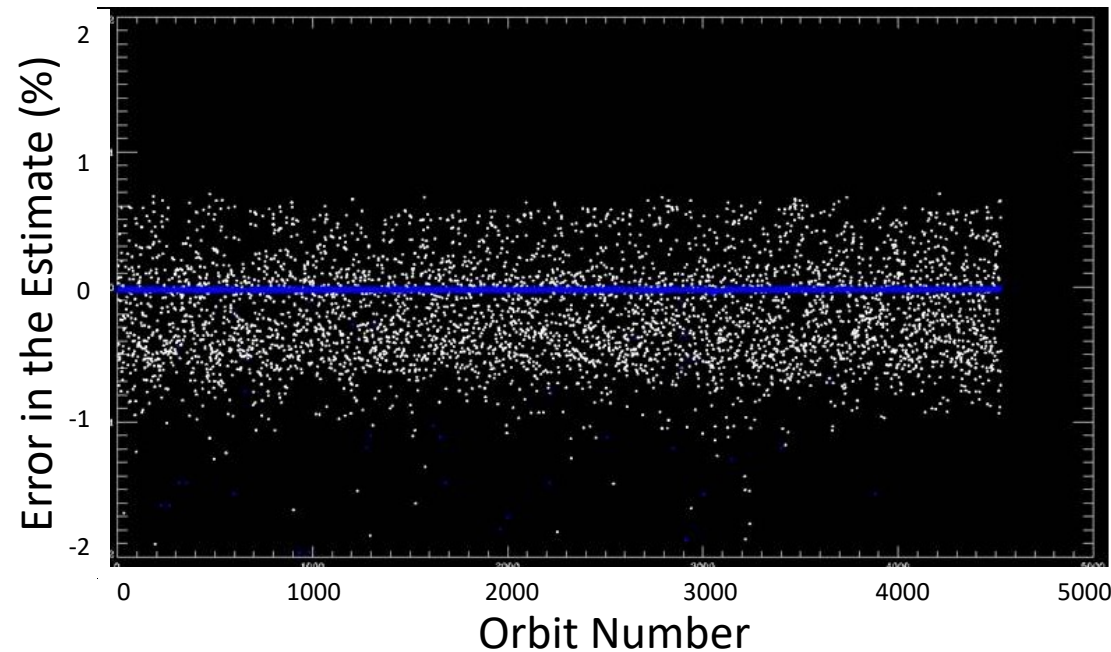
Lunar Observation Angular Sampling Study

Study to quantify the impact of sampling pattern on estimating equivalent full field lunar radiance



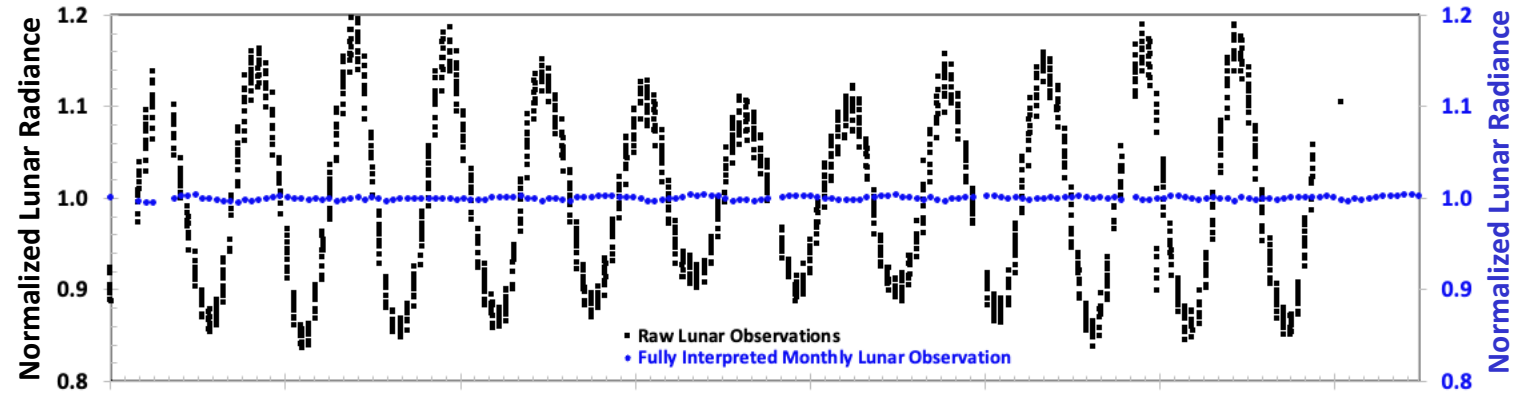
Error in the process of
Estimating equivalent Full
Field Observed Radiance from
a sub Full Field Raster Scan

$\sim \pm 0.3\%$ ($k = 1$)

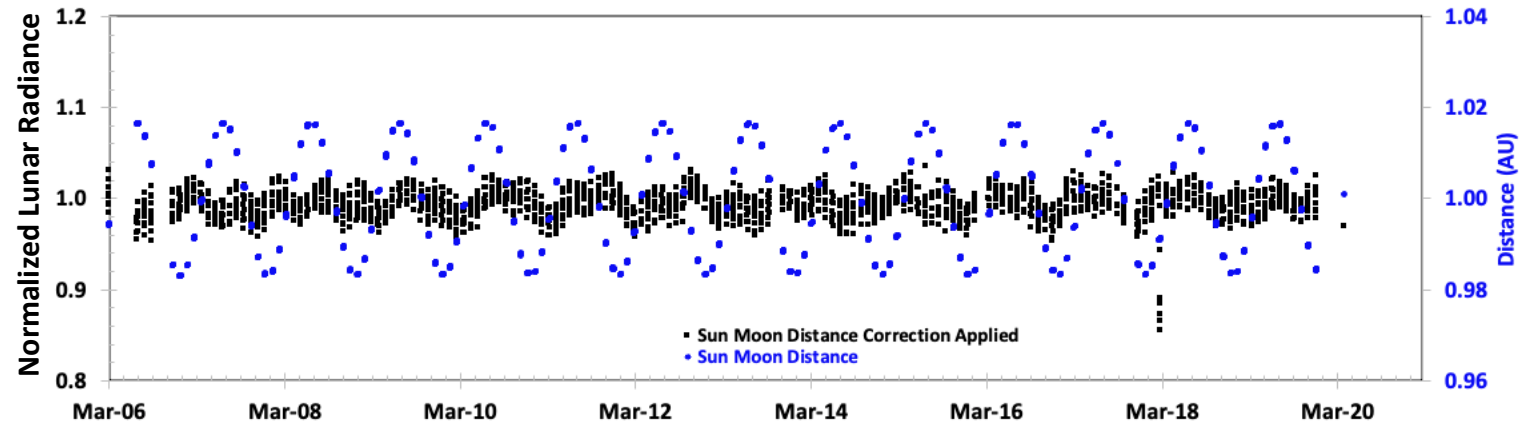
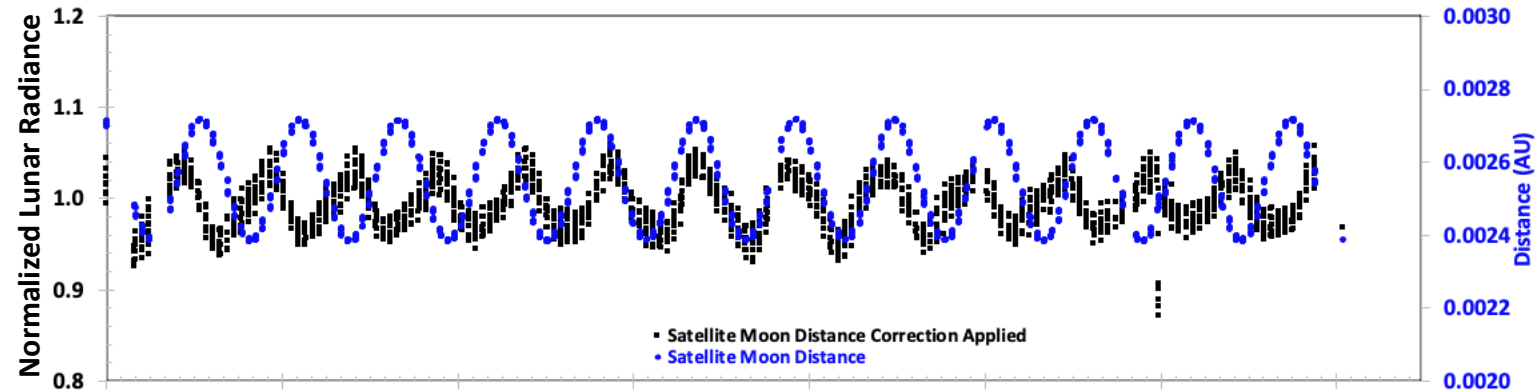


FM1 Total Channel : 2006 to 2021

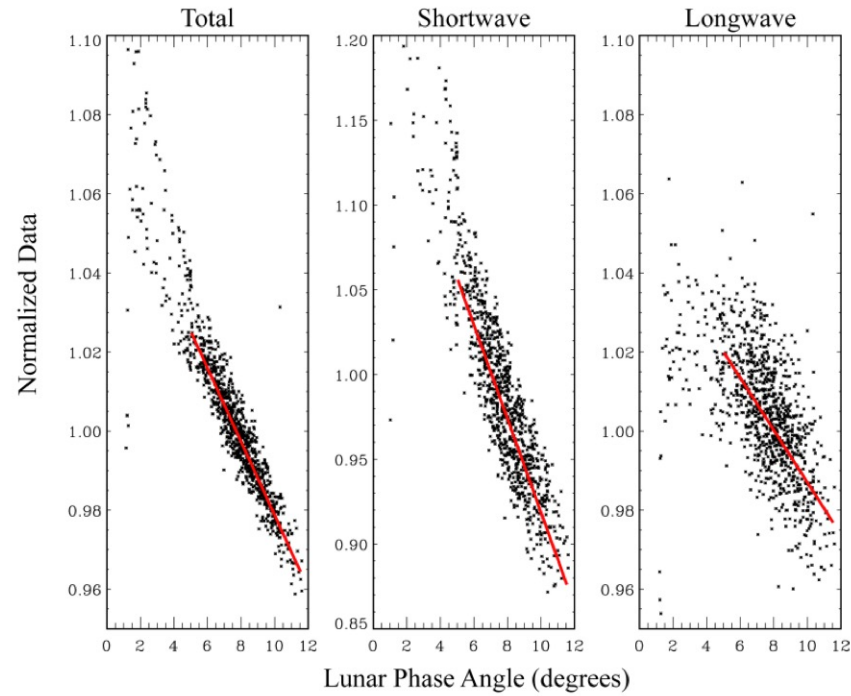
Moon presents
itself to CERES as
a time varying
target of $\sim\pm 20\%$



Once Fully
Interpreted
Variability is at
 $\sim\pm 1\%$

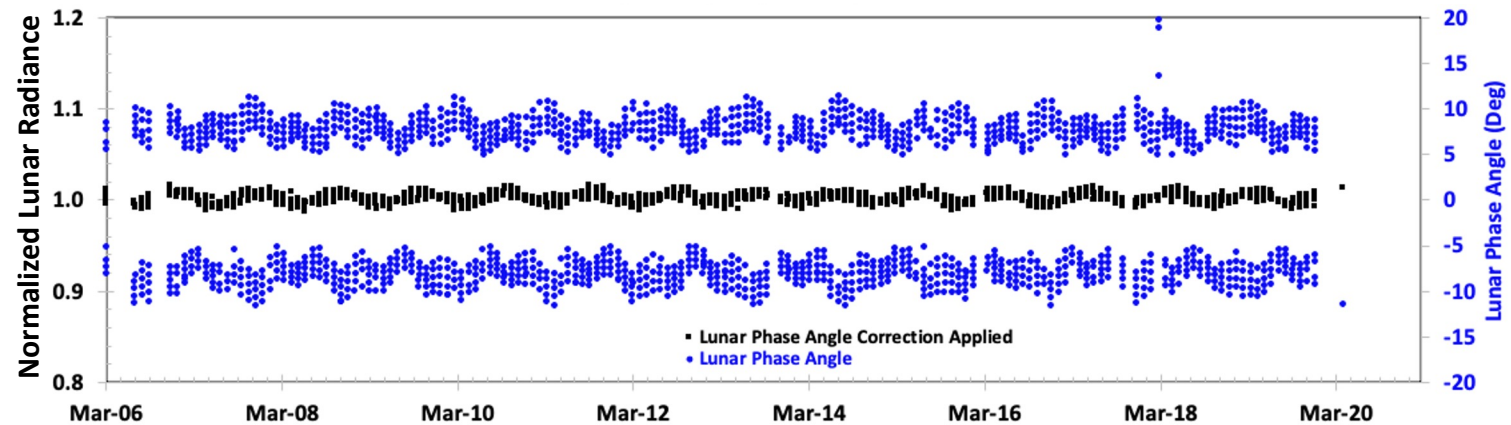


FM1 Total Channel : 2006 to 2021

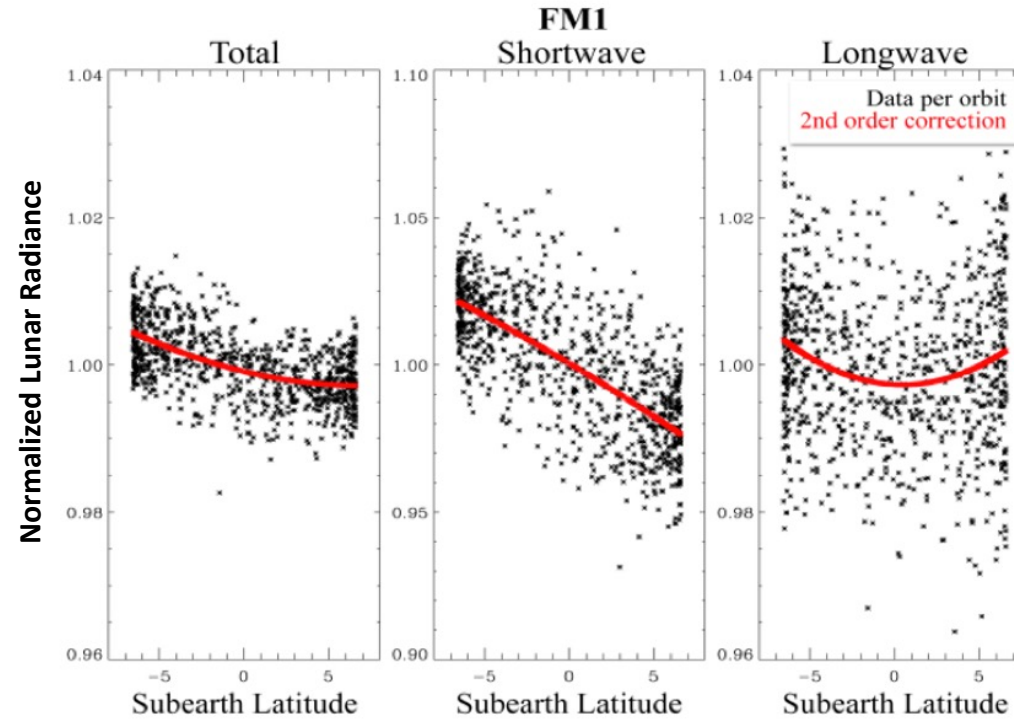


Shortwave Channel
Sensitivity to Lunar
Phase Angle 2X
Total and Longwave

Correct Lunar
Observations to
an 8-deg Lunar
Phase Angle



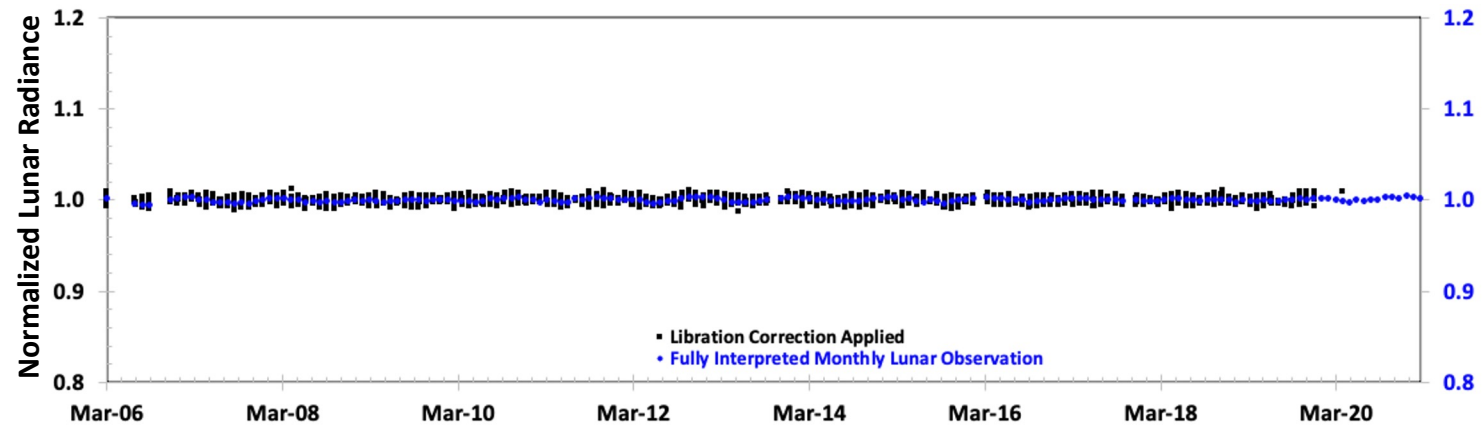
FM1 Total Channel : 2006 to 2021



Shortwave Channel
Sensitivity to Lunar
Libration 2X Total
and Longwave

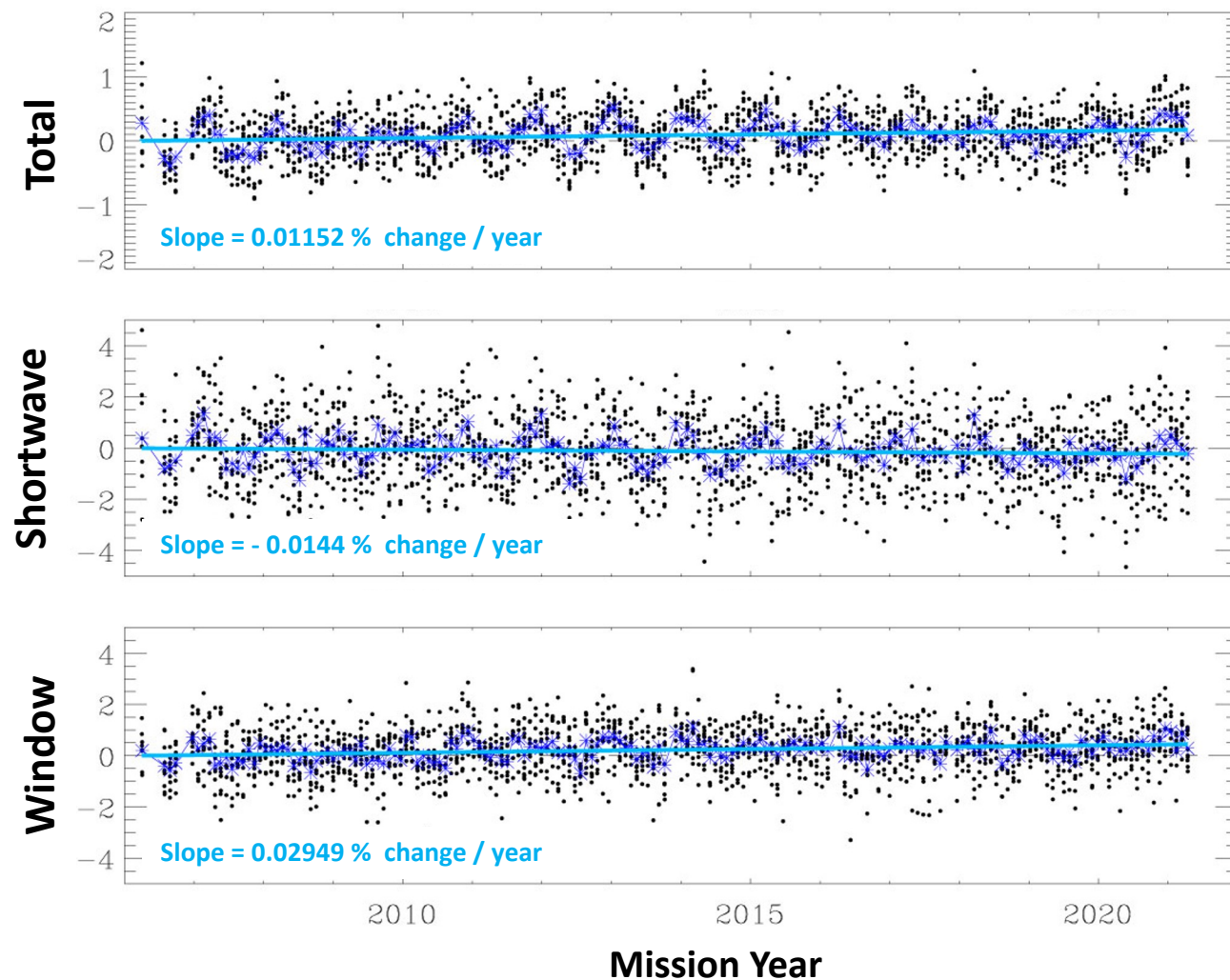
Note: Longitude Correction
Not Shown

Correct Lunar
Observations to
an 8-deg Lunar
Phase Angle

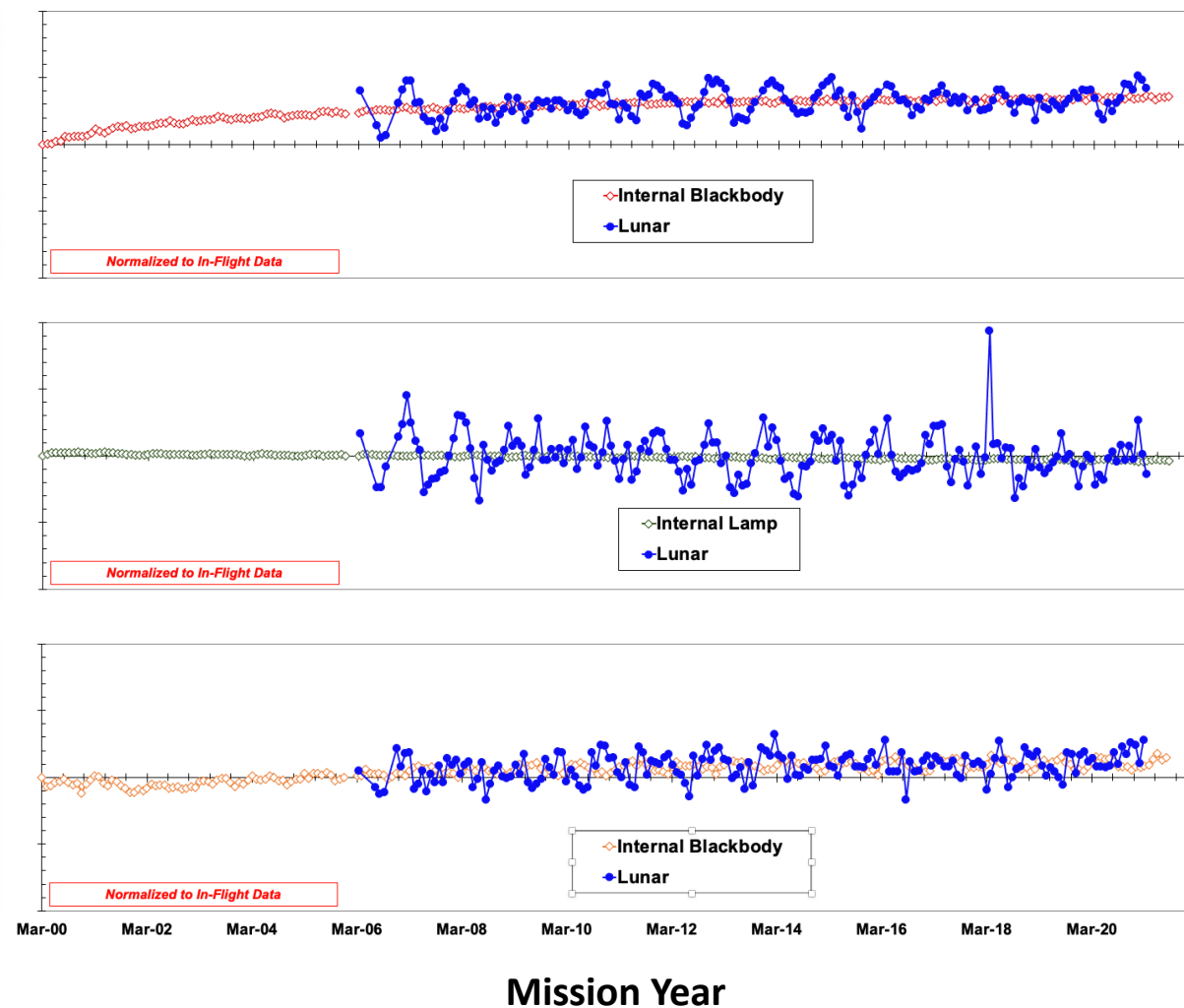


FM-1 Lunar and Internal Calibration Normalized Trend Comparison (%)

Orbital and Lunar Monthly



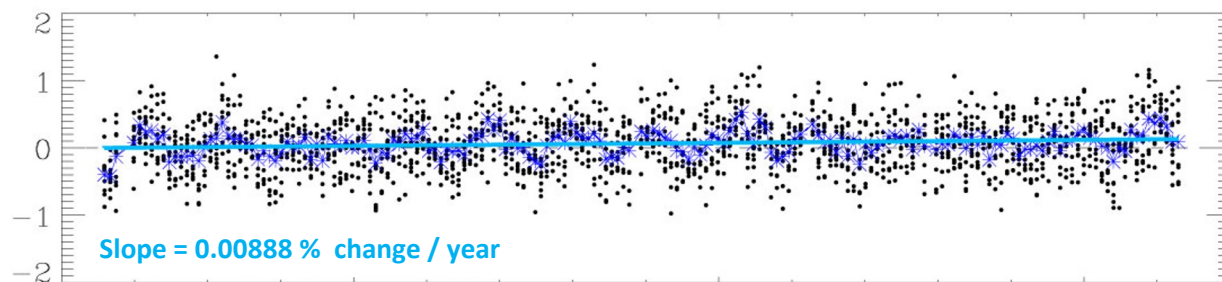
Internal and Lunar Calibrations : Monthly



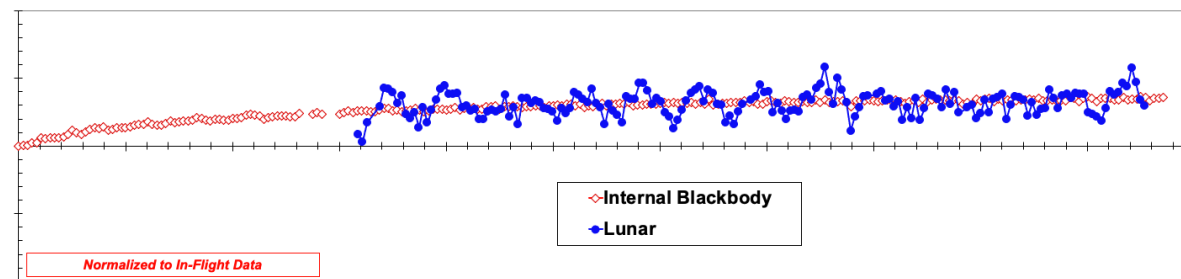
Results display instrument artifacts that have been removed in the Edition-4 Products

FM-2 Lunar and Internal Calibration Normalized Trend Comparison (%)

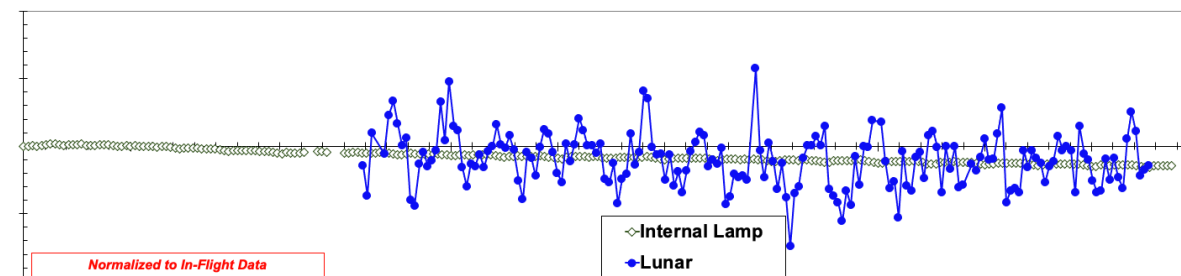
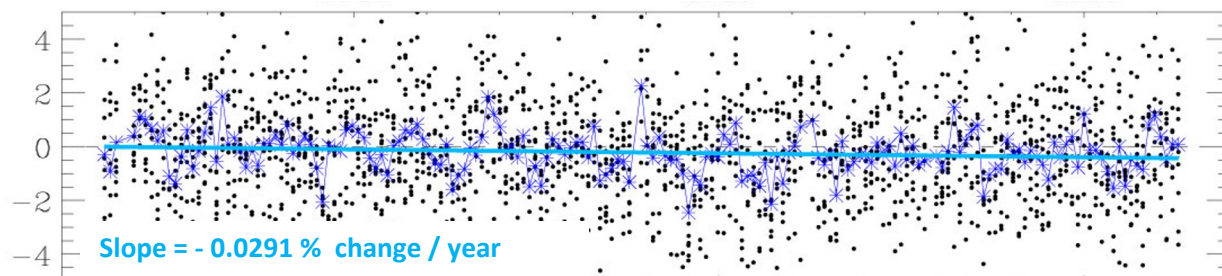
Orbital and Lunar Monthly



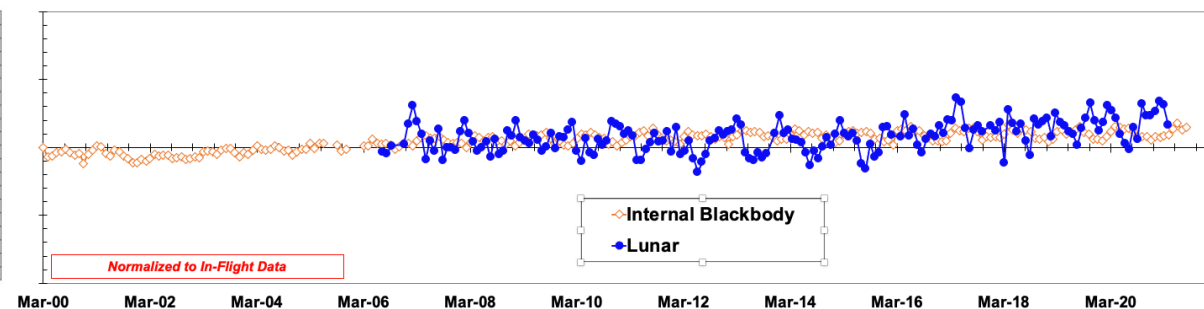
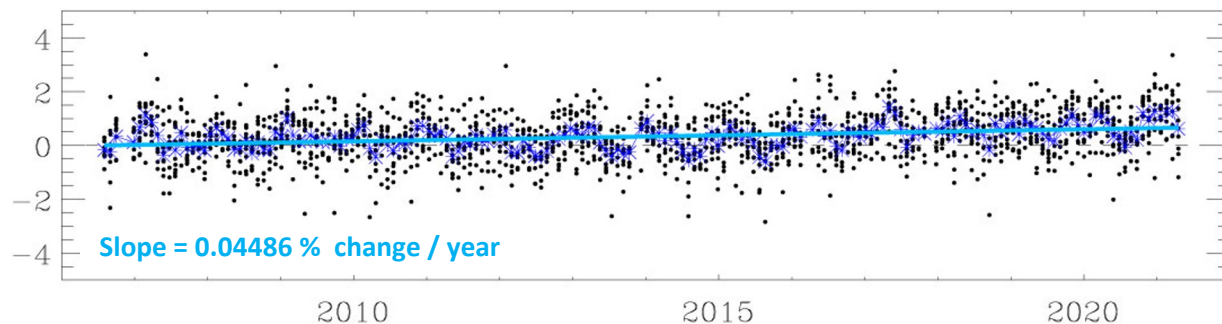
Internal and Lunar Calibrations : Monthly



Shortwave



Window



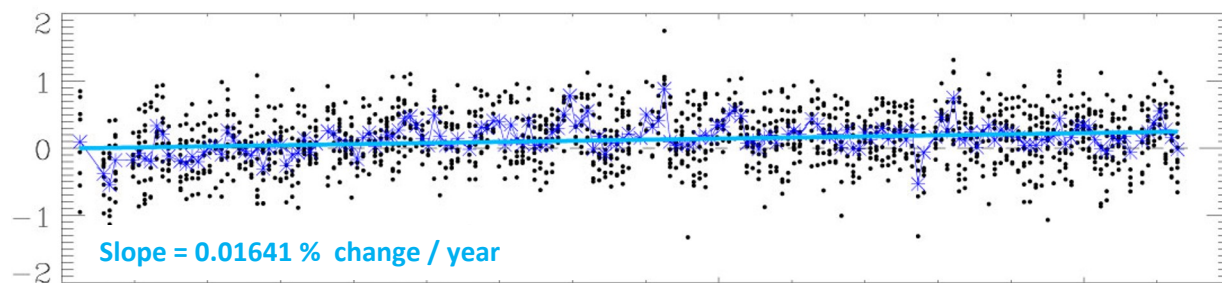
Mission Year

Mission Year

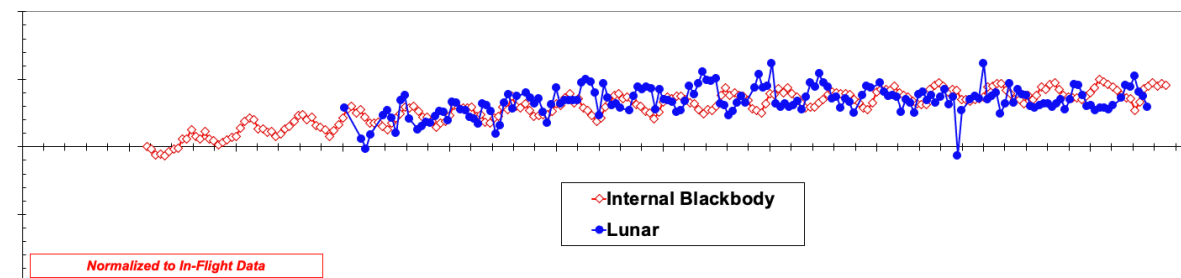
Results display instrument artifacts that have been removed in the Edition-4 Products

FM-3 Lunar and Internal Calibration Normalized Trend Comparison (%)

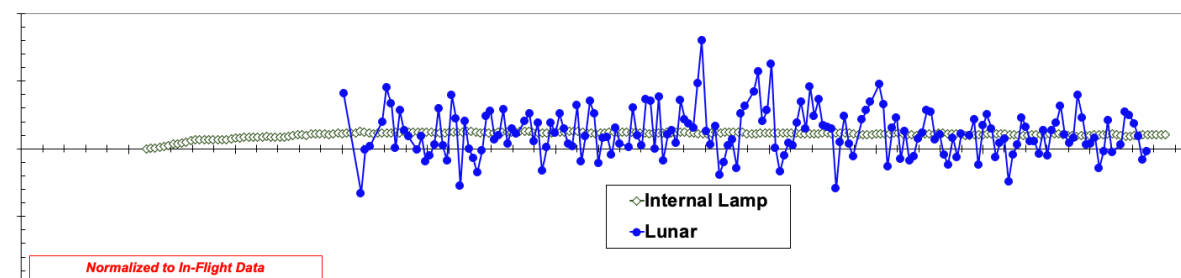
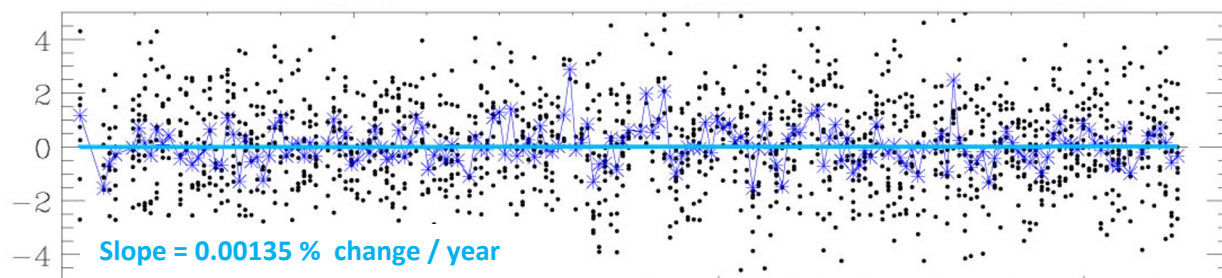
Orbital and Lunar Monthly



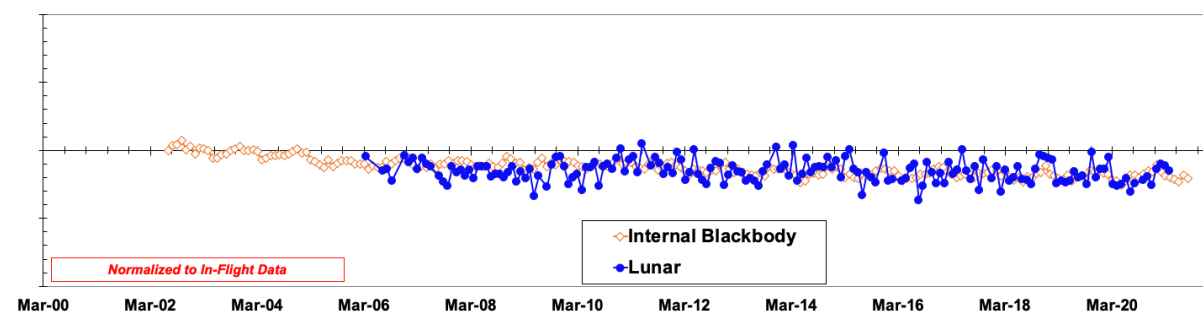
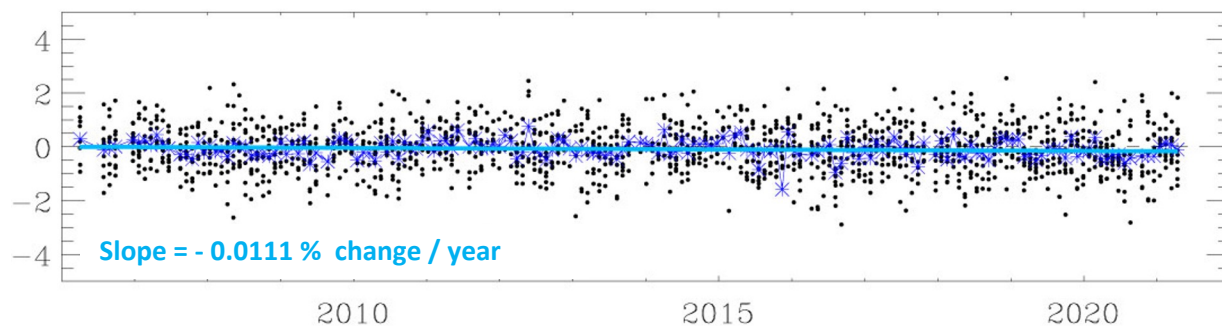
Internal and Lunar Calibrations : Monthly



Shortwave



Window



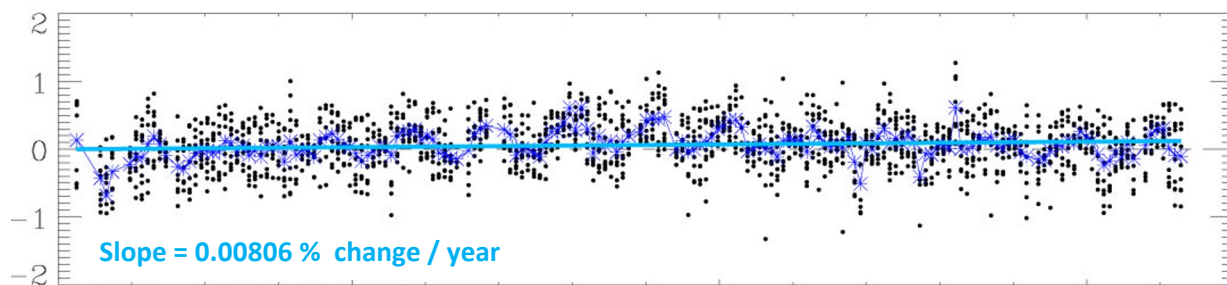
Mission Year

Mission Year

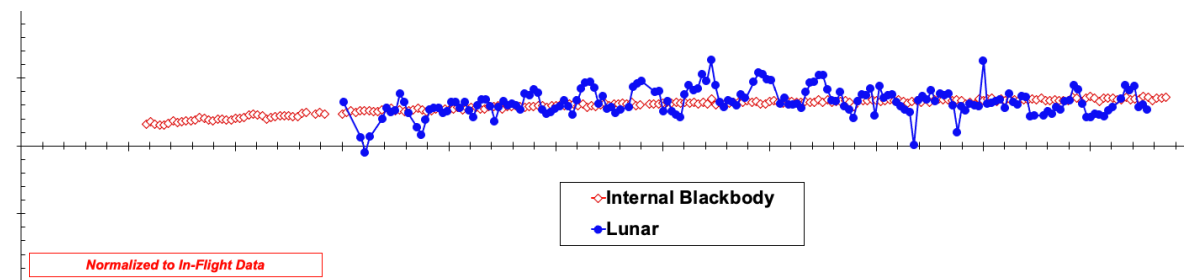
Results display instrument artifacts that have been removed in the Edition-4 Products

FM-4 Lunar and Internal Calibration Normalized Trend Comparison (%)

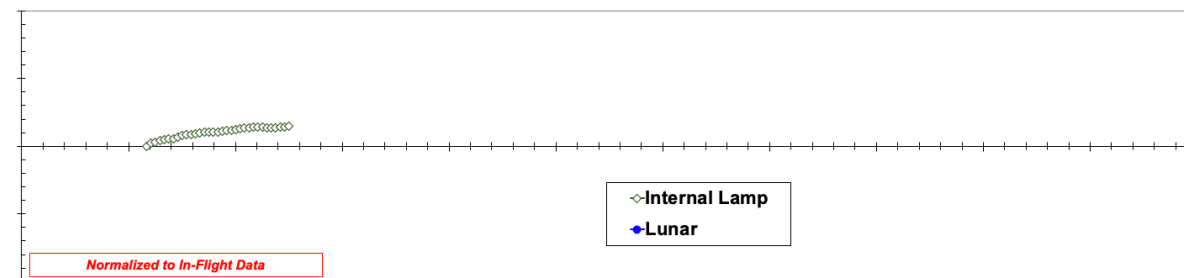
Orbital and Lunar Monthly



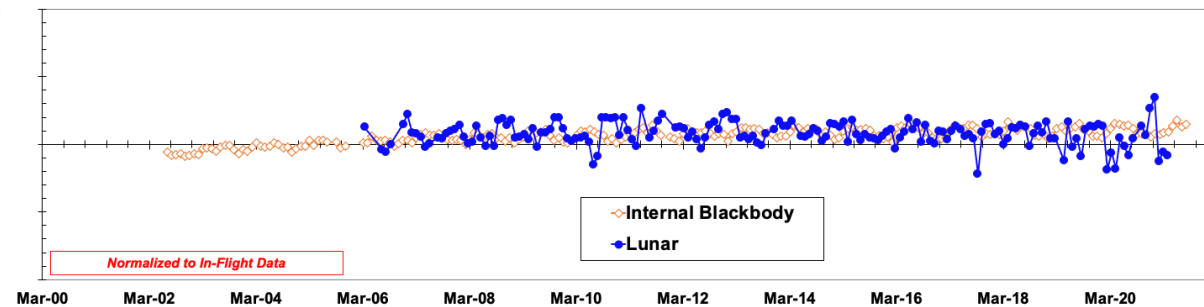
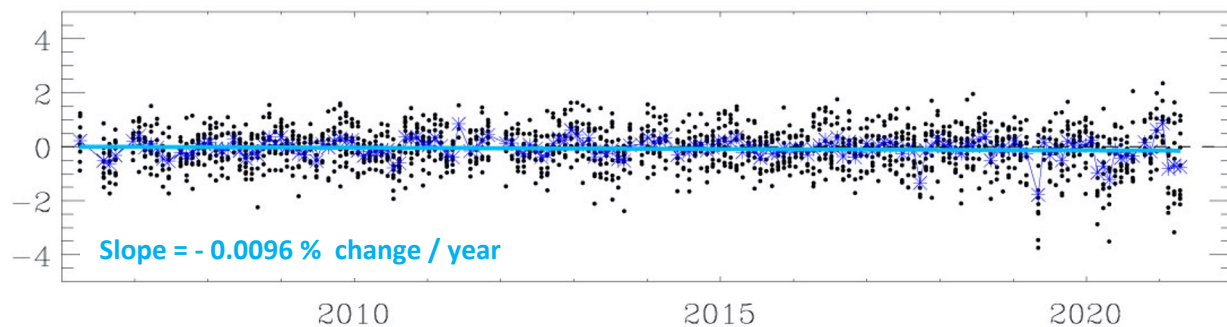
Internal and Lunar Calibrations : Monthly



Shortwave



Window



Mission Year

Mission Year

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- 2) Co-alignment of the three science channels**
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Back-up

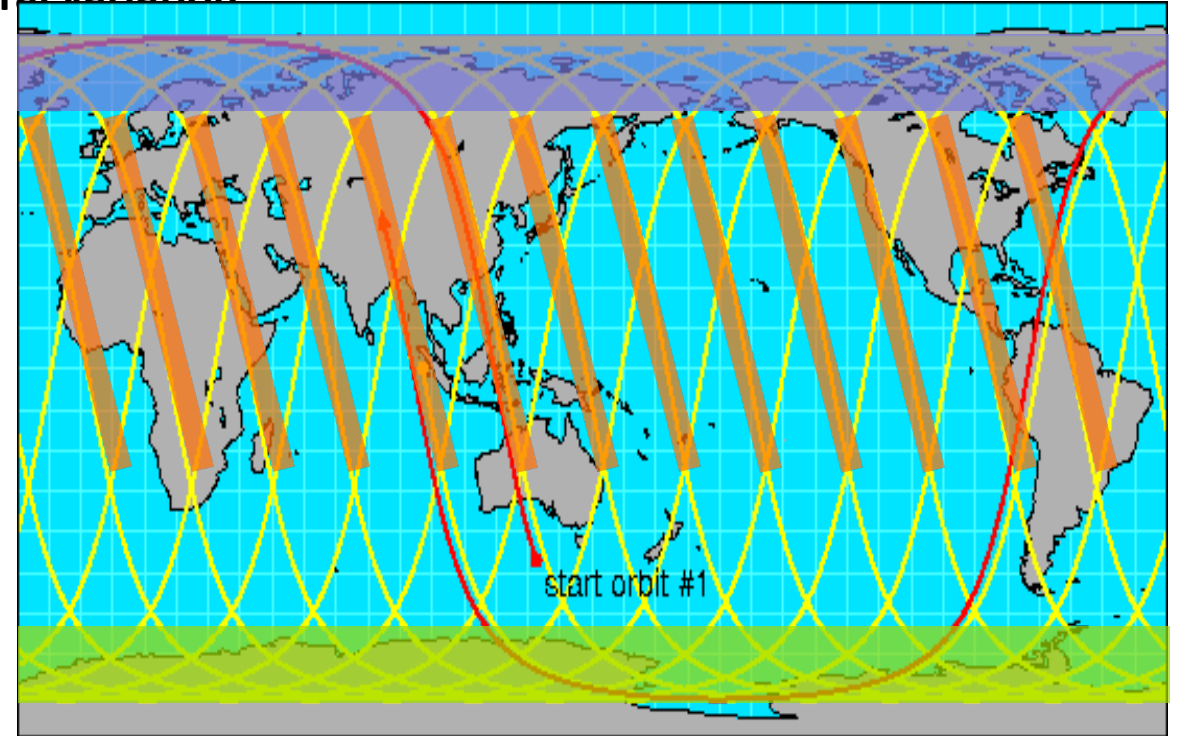
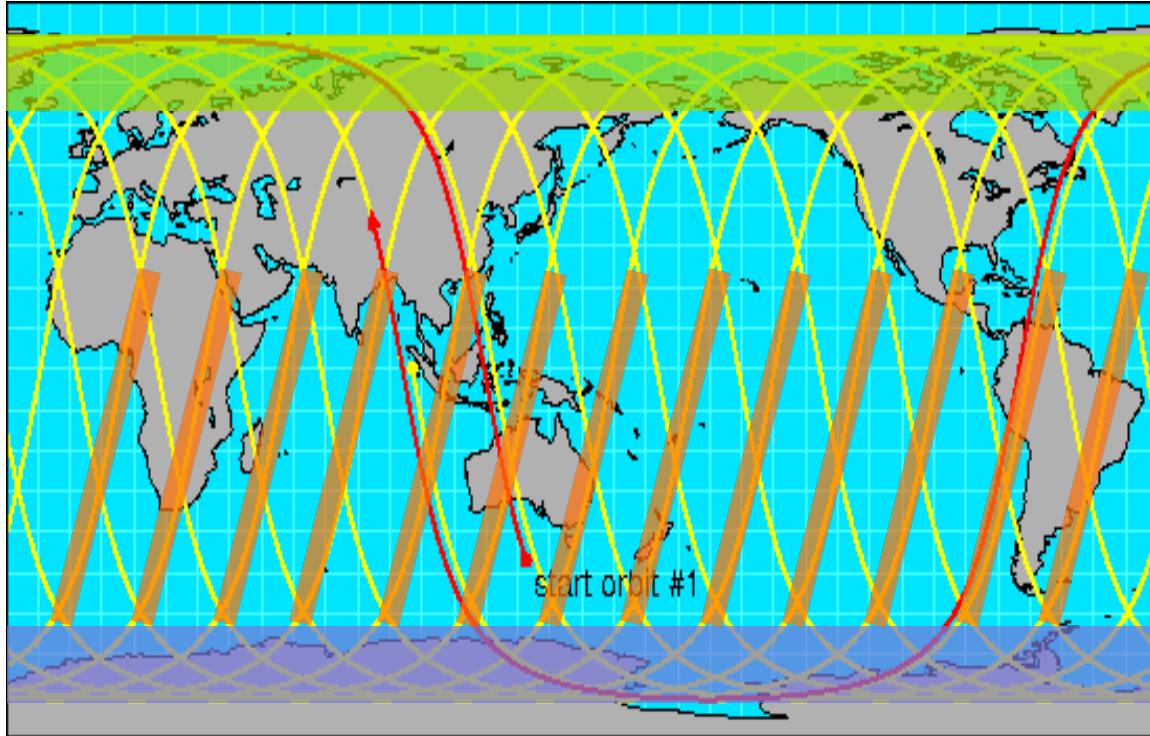
Flight Cal/Val Locations

Still not corrected for:

Aqua

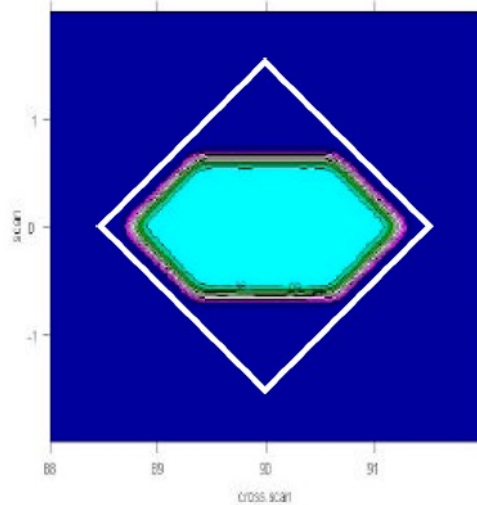
Terra

Spectral variation

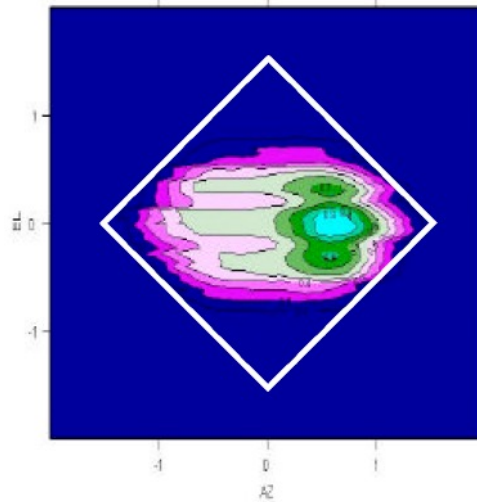


- Lunar Observations
- Solar Calibrations
- Internal Calibration Sequence

Monte-Carlo Ray Trace



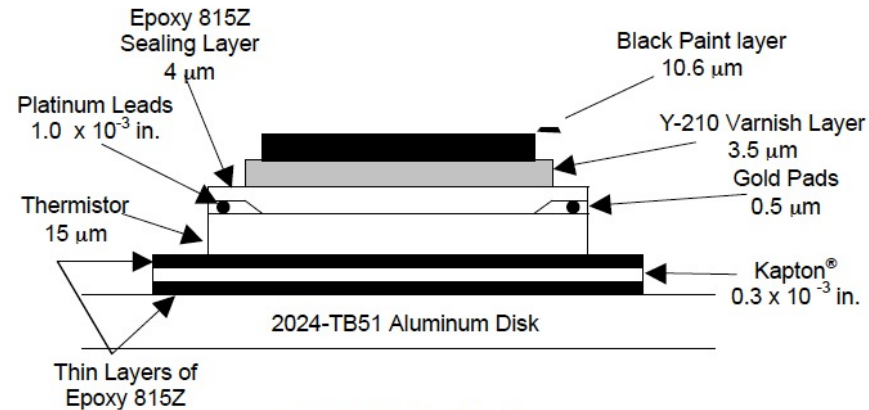
FM-2 WN Channel



FM-2 Window Channel

- One quadrant of detector has an extremely high responsivity
- Probably due to localized delamination of detector due to air-bubbles in the lower Epoxy layers.
 - Consistent with known fabrication problems at the time these detectors were made.
- The resulting void would provide a higher local thermal impedance
 - Inducing higher temperatures in thermistor layer
 - Result in a slower time constant
 - Much longer air-to-vacuum stabilization time

Very stable radiometric performance on orbit!!!!



CERES Detector



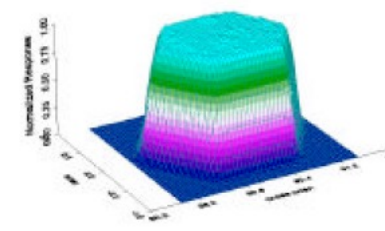
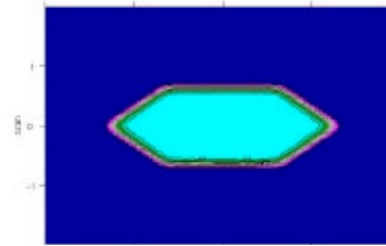
NASA Langley Research Center
Atmospheric Sciences

K. J. Priestley 11/3/98

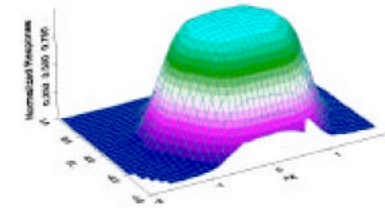
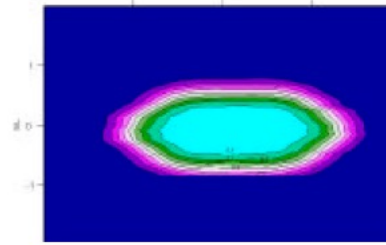


Lunar Scanning Results – CERES Optical Transfer Function

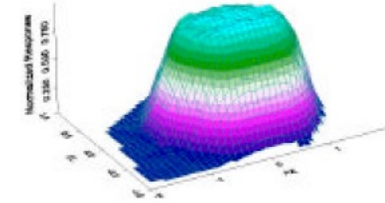
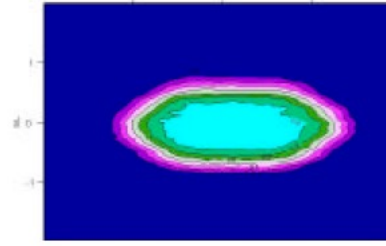
Monte-Carlo Ray Trace
(FELIX)



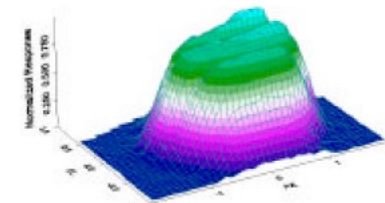
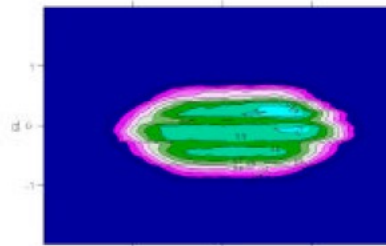
FM-1 Total



FM-1 Shortwave



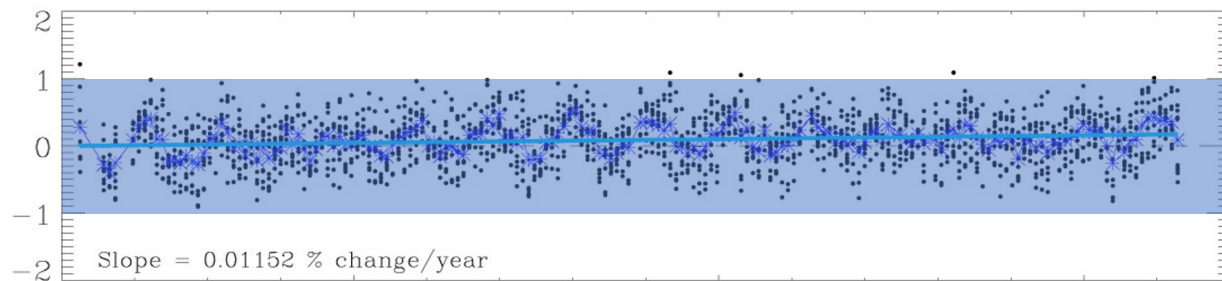
FM-1 Window



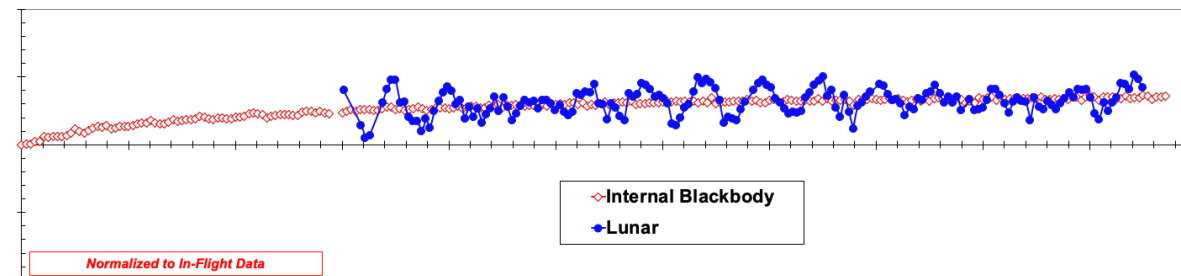
- No longer have working group meetings at Science Team so thought we would initiate individual focused presentations on each of the cal val activities
- Recognize arcstone selection
- Provide inputs to costly and Cindy on CERES perspective
- Timeline of when and why we executed lunar scanning
- Wide field telescope introduces it's own uniqueness vs imagers with much smaller IFOV
- Phases of experiment
- Operations to sample
- Analysis of collected data (Arcstone may help here)
- Comparison to other cal/val experiment results
- Lunar reflectance plot
-
- Time scales.... Lunar is one data point per month
- See calcon plot for where calibrations occur
- The moon is okay for what it is, but it's really only one element in an overall cal/val protocol

FM1 Lunar and Internal Calibration Trend Comparison (%)

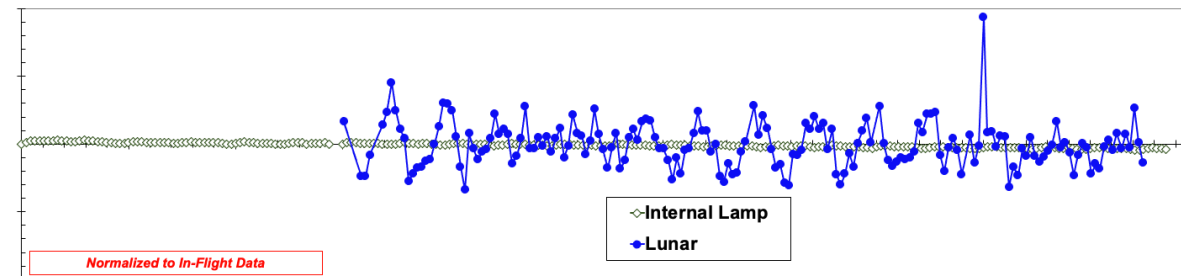
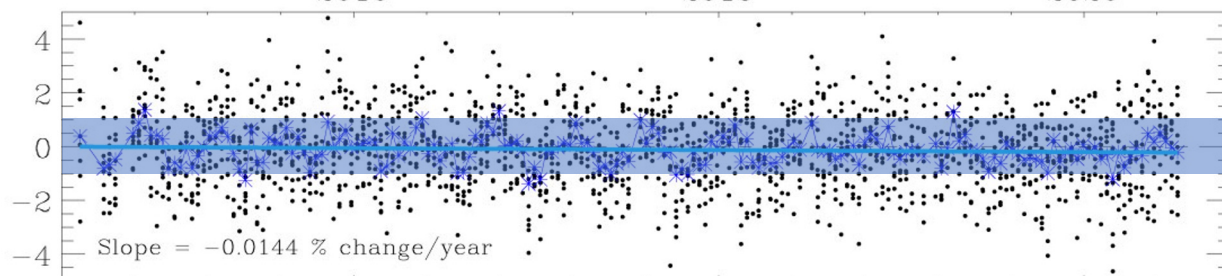
Orbital and Lunar Monthly



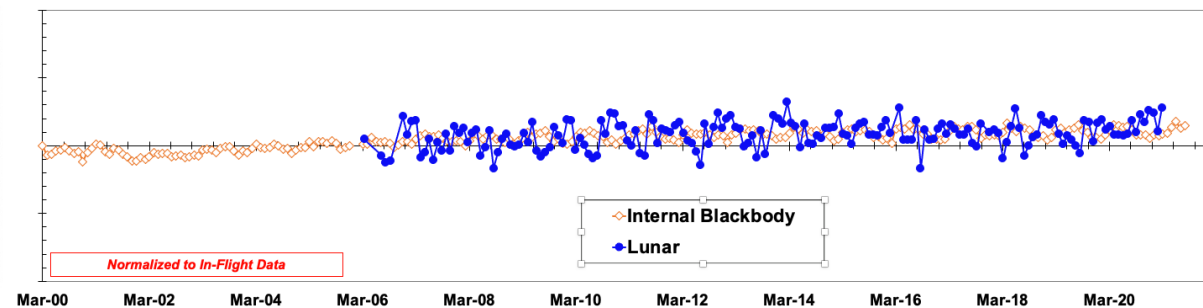
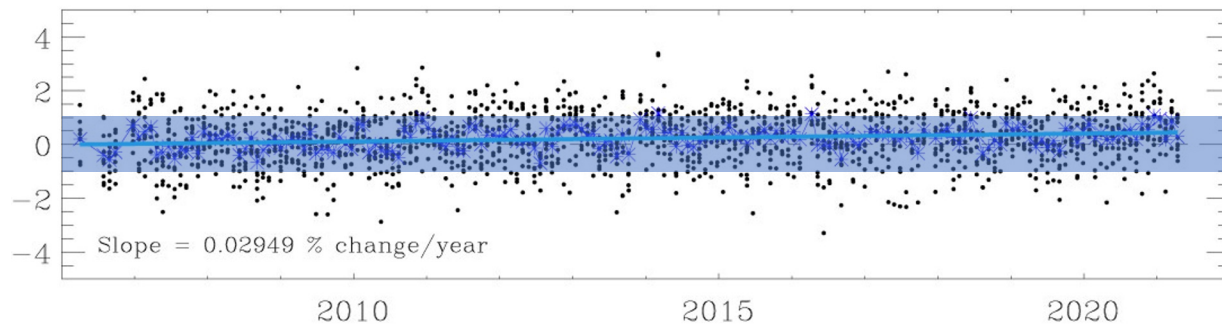
Internal and Lunar Calibrations : Monthly



Shortwave



Window

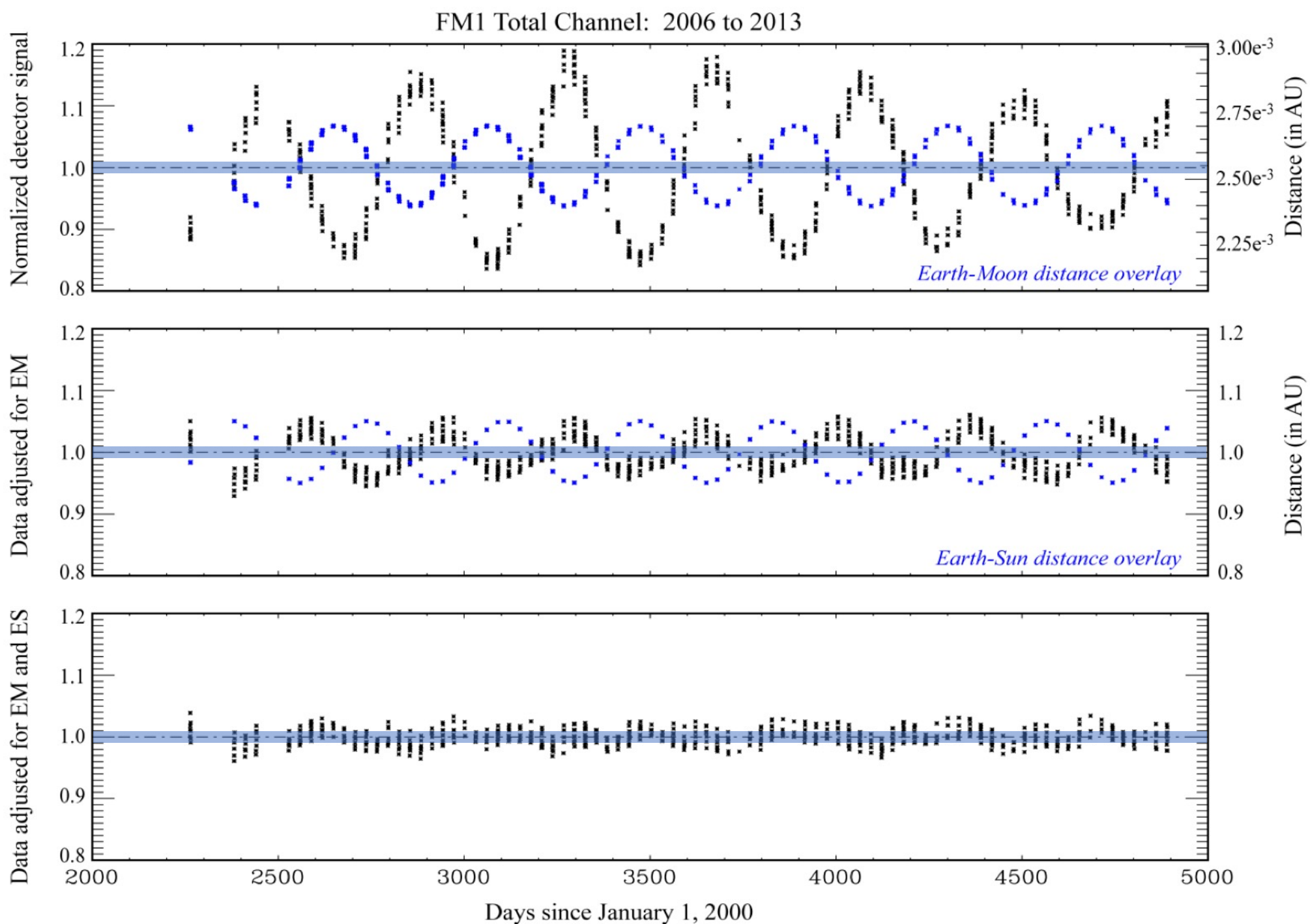


Mission Year

Mission Year

Moon presents itself to CERES as a time varying target of $\sim\pm 20\%$

To be useful peak to peak variability/noise must be reduced to better than $\pm 1\%$.



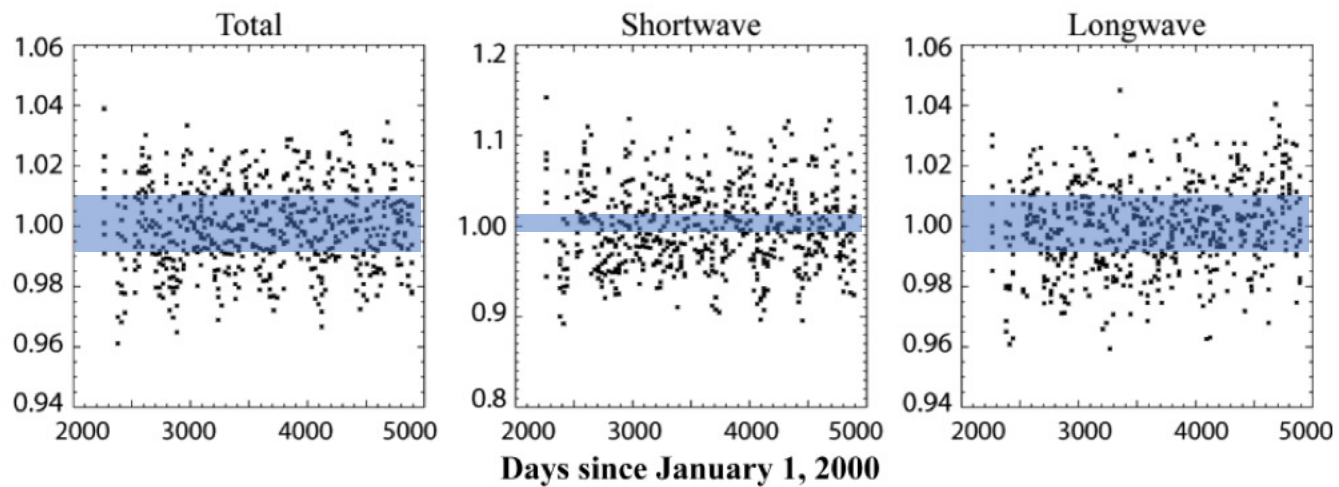


Fig. 11. CERES FM-1 detector response before lunar phase angle adjustment.

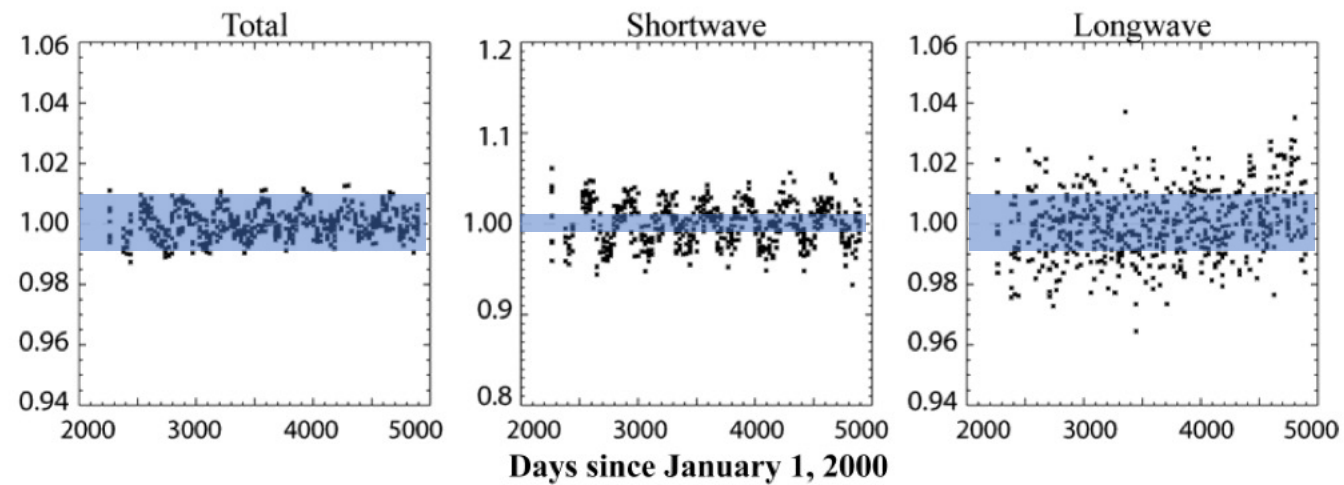


Fig. 12. CERES FM-1 detector response after lunar phase angle adjustment.

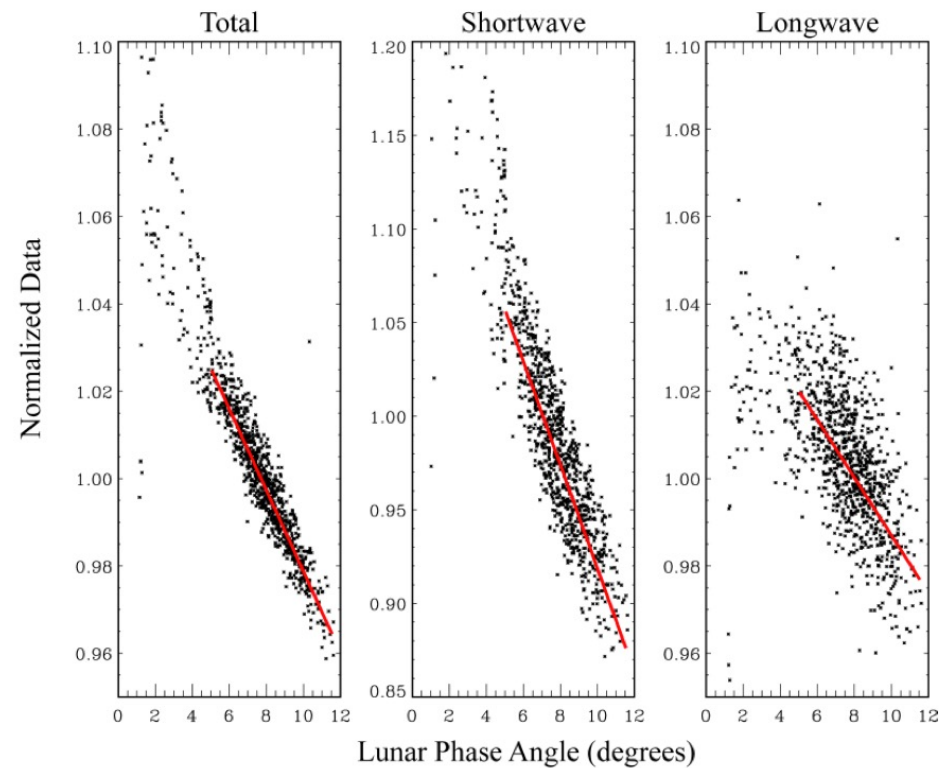


Fig. 10. Lunar Phase Angle and CERES detector response

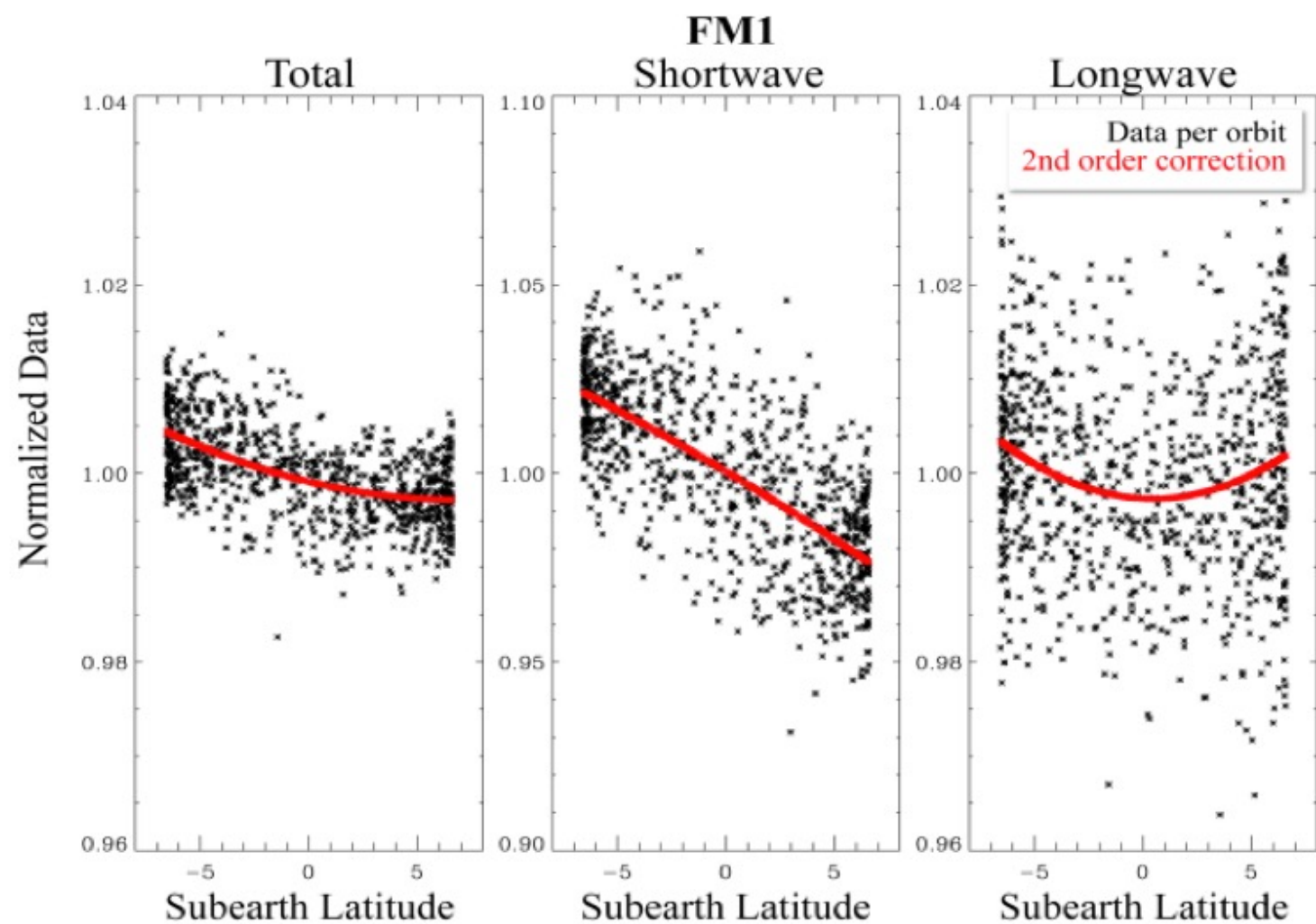


Fig. 14. FM-1 detector output versus lunar libration latitude.

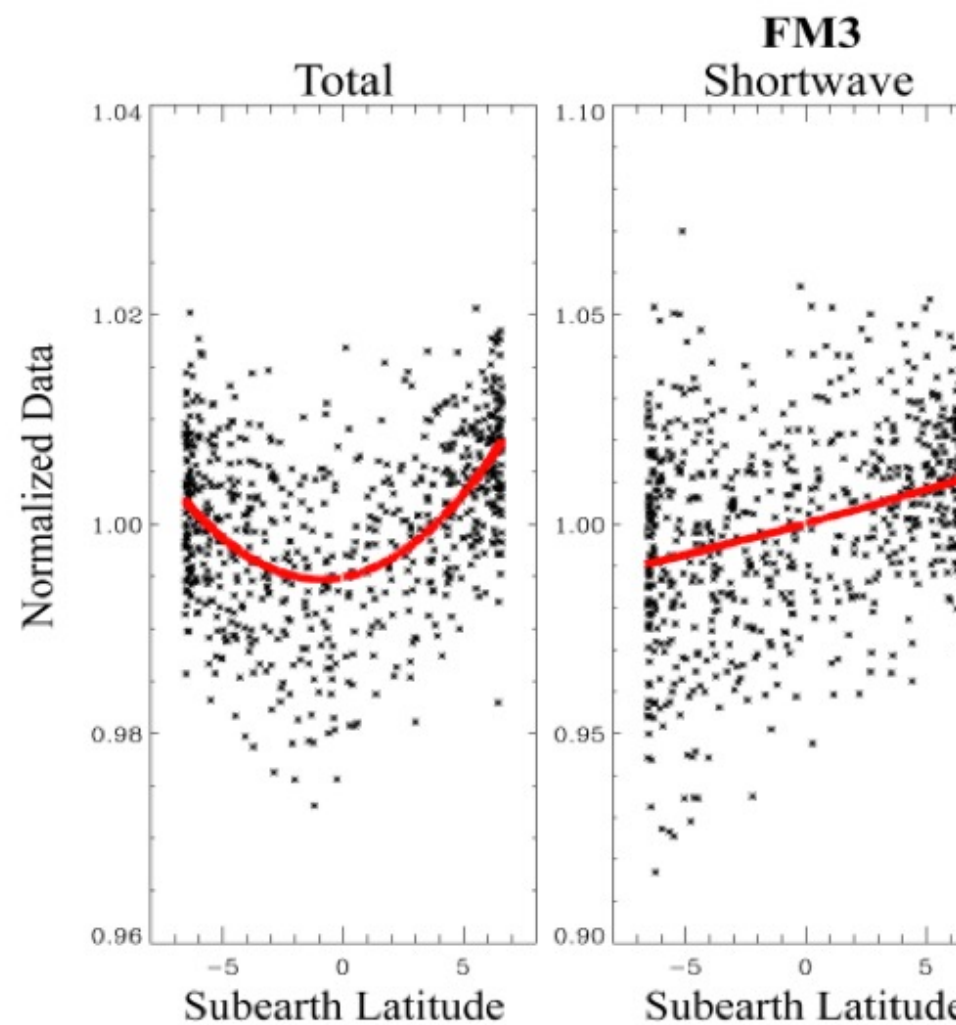


Fig. 15. FM-3 detector output versus lunar libration latitude.